Evidence for spatial variation of $\boldsymbol{\alpha}$

 Webb, King, Murphy, Flambaum, Carswell, Bainbridge,arxiv:1008.3907,
 α(x)= α(0) + α '(0) x + ... x=r cos(φ), r=ct – distance (instead of time)

Reconciles all measurements of the variation

- Berengut, Flambaum, arxiv:1008.3957
- Manifestations on atomic clocks, Oklo, meteorites and cosmological phenomena
- Berengut, Flambaum,King,Curran,Webb,
 Further evidence

Variation of Fundamental Constants

V.V. Flambaum School of Physics, UNSW, Sydney, Australia Co-authors: Atomic calculations V.Dzuba, M.Kozlov, E.Angstmann, J.Berengut,M.Marchenko,Cheng Chin,S.Karshenboim,A.Nevsky, S.Porsev Nuclear and QCD calculations E.Shuryak, V.Dmitriev, D.Leinweber, A.Thomas, R.Young, A.Hoell, P.Jaikumar, C.Roberts,S.Wright, A.Tedesco, W.Wiringa Cosmology J.Barrow Quasar data J.Webb,M.Murphy, J.King, S.Curran, M.Drinkwater,P.Tsanavaris, C.Churchill,J.Prochazka,A.Wolfe,S.Muller,C,Henkel, F.Combes, T.Wiklind, R.Carswell,M.Bainbridge Laboratory measurements S.J. Ferrel,A,Cingoz,ALappiere,A.-T.Nguyen,N.Leefer, D.Budker,S.K.Lamoreuax,J.R.Torgerson,S.Blatt,A.D.Ludlow,G.K.Cambell, J.W.Thomsen,T.Zelevinsky,M.M.Boid,J.Ye,X.Baillard,M.Fouche,R.LeTargat,A.Brush, P.Lemonde,M.Takamoto,F.-L.Hong,H.Katori

Motivation

- Extra space dimensions (Kaluza-Klein, Superstring and M-theories). Extra space dimensions is a common feature of theories unifying gravity with other interactions. Any change in size of these dimensions would manifest itself in the 3D world as variation of fundamental constants.
- Scalar fields . Fundamental constants depend on scalar fields which vary in space and time (variable vacuum dielectric constant ε_0). May be related to "dark energy" and accelerated expansion of the Universe..
- "Fine tuning" of fundamental constants is needed for humans to exist. Example: low-energy resonance in production of carbon from helium in stars (He+He+He=C). Slightly different coupling constants — no resonance — no life.

Variation of coupling constants in space provide natural explanation of the "fine tuning": we appeared in area of the Universe where values of fundamental constants are suitable for our existence.

Search for variation of fundamental constants

•Big Bang Nucleosynthesis

•Quasar Absorption Spectra ¹

Oklo natural nuclear reactor

•Atomic clocks ¹

Enhanced effects in atoms ¹, molecules¹ and nuclei
Dependence on gravity

¹ Based on atomic and molecular calculations

evidence? evidences?

Dimensionless Constants

- Since variation of <u>dimensional</u> constants cannot be distinguished from variation of <u>units</u>, it only makes sense to consider variation of <u>dimensionless</u> constants.
- Fine structure constant $\alpha = e^2/2\varepsilon_0 hc = 1/137.036$
- Electron or quark mass/QCD strong interaction scale, $m_{e,q}/\Lambda_{QCD}$ α_{strong} (r)=const/ln(r Λ_{QCD} /ch)

Variation of strong interaction

Grand unification

$$\frac{\Delta \left(m / \Lambda_{QCD} \right)}{m / \Lambda_{QCD}} = R \frac{\Delta \alpha}{\alpha}$$

- 1. Proton mass $M_p = 3\Lambda_{QCD}$, measure m_e / M_p
- 2. Nuclear magnetic moments

$$\mu = g e \hbar / 4M_p c, \quad g = g \left(m_q / \Lambda_{QCD} \right)$$

3. Nuclear energy levels and resonances

Variation of strong interaction

Grand unification (Calmet, Fritzsch; Langecker, Segre, Strasser; Wetterich, Dent)

$$\frac{\Delta \left(m / \Lambda_{QCD} \right)}{m / \Lambda_{QCD}} \quad 35 \frac{\Delta \alpha}{\alpha}$$

- 1. Proton mass $M_p = 3\Lambda_{QCD}$, measure m_e / M_p
- 2. Nuclear magnetic moments

$$\mu = g e \hbar / 4M_p c, \quad g = g \left(m_q / \Lambda_{QCD} \right)$$

3. Nuclear energy levels and resonances

Relation between variations of different coupling constants Grand unification models Calmet,Fritzch; Langecker, Segre, Strasser; Wetterich,Dent

 $\alpha_i^{-1}(v) = \alpha_{GUT}^{-1} + b_i \ln(v / v_0)$ Variation of GUT const α_{GUT} $d\alpha_1^{-1} = d\alpha_2^{-1} = d\alpha_3^{-1} = d\alpha_{GUT}^{-1}$ $d\alpha_3 / \alpha_3^{-2} = d\alpha_1 / \alpha_1^{-2}$

$$\alpha_{3}^{-1}(m) = \alpha_{strong}^{-1}(m) = b_{3}ln(m / \Lambda_{QCD})$$

$$\alpha^{-1}(m) = 5/3 \alpha_{1}^{-1}(m) + \alpha_{2}^{-1}(m)$$

$$\frac{\Delta(m / \Lambda_{QCD})}{m / \Lambda_{QCD}} = \frac{1}{b_{3}\alpha_{3}} \frac{\Delta \alpha_{3}}{\alpha_{3}} = \frac{const}{\alpha} \frac{\Delta \alpha}{\alpha} \quad 35 \frac{\Delta \alpha}{\alpha}$$
Proton mass M_p = 4 Λ_{QCD} , measure m_{e} / M_{p}

2. Nuclear magnetic moments

1.

$$\mu = g e \hbar / 4M_p c, \quad g = g \left(m_q / \Lambda_{QCD} \right)$$

3. Nuclear energy levels and resonances

Dependence on quark mass

- Dimensionless parameter is m_q/Λ_{QCD} . It is convenient to assume Λ_{QCD} =const, i.e. measure m_q in units of Λ_{QCD}
- m_{π} is proportional to $(m_q\Lambda_{QCD})^{1/2}$ $\Delta m_{\pi}/m_{\pi}{=}0.5\Delta m_q/m_q$
- Other meson and nucleon masses remains finite for $m_q=0$. $\Delta m/m=K \Delta m_q/m_q$

Argonne: K are calculated for $p, n, \rho, \omega, \sigma$.

$$m_q = \frac{m_u + m_d}{2} \approx 4 MeV, \Lambda_{QCD} = 220 MeV \rightarrow K = 0.02 - 0.06$$

Strange quark mass $m_s = 120 MeV$

Nuclear magnetic moments depends on π -meson mass m $_{\pi}$





Spin-spin interaction between valence and core nucleons

Nucleon magnetic moment $\mu = \mu_0(1 + am_{\pi} + ...) = \mu_0(1 + b\sqrt{m_q} + ...)$ Nucleon and meson masses

 $M = M_0 + am_q$ QCD calculations: lattice, chiral perturbation theory, cloudy bag model, Dyson-Schwinger and Faddeev equations, semiempirical. Nuclear calculations: meson exchange theory

of strong interaction. Nucleon mass in kinetic

energy p²/2M

Big Bang nucleosynthesis: dependence on quark mass

- Flambaum, Shuryak 2002
- Flambaum, Shuryak 2003
- Dmitriev, Flambaum 2003
- Dmitriev, Flambaum, Webb 2004
- Coc, Nunes, Olive, Uzan, Vangioni 2007
- Dent, Stern, Wetterich 2007
- Flambaum, Wiringa 2007
- Berengut, Dmitriev, Flambaum 2009

Big Bang Nucleosynthesis (Dmitriev, Flambaum, Webb)



 $p + n \rightarrow d + \gamma$, $3 \sec \le t \le 6 \min$

Productions of D, ⁴He, ⁷Li are exponentially sensitive to deuteron binding energy E_d

i
$$\sim e^{-\frac{E_d}{T_f}}$$

- η from cosmic microwave background fluctuations (η - barion to photon ratio).

- η from BBN for present value of Q $(Q = |E_d|)$



FIG. 2.—Evolution of light-element abundances with temperature, for a baryon-to-photon ratio $\eta_{10} = 3.16$. The dashed curves give the NSE curves of ⁴He, t, ³He, and d, respectively. The dotted curve is explained in the text.

Deuterium bottleneck

- At temeperature T<0.3 Mev all abundances follow deuteron abundance
- (no other nuclei produced if there are no deuterons)
 - Reaction γ d n p , exponentially small number of energetic photons, e^{-(Ed/T)}
- Exponetilal sensitivity to deuteron binding energy E_d , E_d =2 Mev,
- Freezeout temeperure $T_f = 30 \text{ KeV}$



Flambaum, Shuryak: Deuteron Binding Energy is very sensitive to variation of *strange* quark mass (4 factors of enhancement):

1. Deuteron is a shallow bound level.

Virtual level in $n+p \rightarrow d+\gamma$ is even more sensitive to the variation of the potential.



2. Strong compensation between σ -meson and ω -meson exchange in potential (Walecka model): $4\pi rV = -g_s^2 e^{-m_\sigma r} + g_v^2 e^{-m_\omega r}$

3. $\sigma = \frac{1}{\sqrt{3}}(u\overline{u} + d\overline{d} + s\overline{s}), \quad m_{\sigma} \approx \frac{2}{3}m_s + 2\Lambda_{QCD}$

4. Repulsion of σ from $K\overline{K}$ threshold Total $\frac{\delta E_d}{E_d} \approx -17 \frac{\delta m_s}{m_s}$ and $\frac{\delta (m_s/\Lambda_{QCD})}{m_s/\Lambda_{QCD}} = (+1.1\pm0.3) \times 10^{-3}$

New BBN result

- Dent,Stern,Wetterich 2007; Berengut, Dmitriev, Flambaum 2009: dependence of BBN on energies of ^{2,3}H.^{3,4}He.^{6,7}Li.^{7,8}Be
- Flambaum, Wiringa 2007 : dependence of binding energies of ^{2,3}H,^{3,4}He,^{6,7}Li, ^{7,8}Be on nucleon and meson masses,
- Flambaum,Holl,Jaikumar,Roberts,Write, Maris 2006: dependence of nucleon and meson masses on light quark mass m_q.

Big Bang Nucleosynthesis: Dependence on m_q/Λ_{QCD}

- 2 H 1+7.7x=1.07(15) x=0.009(19)
- ⁴He 1-0.95x=1.005(36) x=-0.005(38)
- ⁷Li 1-50x=0.33(11) x=0.013(02)

Final result

 $x = \Delta X_q / X_q = 0.013 (02), X_q = m_q / \Lambda_{QCD}$

Big Bang Nucleosynthesis: Dependence on m_q/Λ_{QCD}

- ²H 1+7.7x=1.07(15) x=0.009(19)
- ⁴He 1-0.95x=1.005(36) x=-0.005(38)
- ⁷Li 1-50x=0.33(11) x=0.013(02) result

 $x=\Delta X_q/X_q = 0.013 (02), X_q=m_q/\Lambda_{QCD}$ Dominated by ⁷Li abundance (3 times difference), consistent with ²H,⁴He Nonlinear effects: $x=\Delta X_q/X_q = 0.016 (05)$

Atomic transition frequencies

Use atomic calculations to find $\omega(\alpha)$.

For α close to α_0 $\omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$

q is found by varying α in computer codes:

 $q = d\omega/dx = [\omega(0.1) - \omega(-0.1)]/0.2, x = \alpha^2/\alpha_0^2 - 1$

Many-Multiplet Method Dzuba,Flambaum,Webb 1998 quasar spectroscopy and atomic clocks

Variation of fine structure constant $\boldsymbol{\alpha}$

Many-Multiplet Method

Relativistic correction to electron energy E_n :

$$\Delta_n = \frac{E_n}{\nu} (Z\alpha)^2 \left[\frac{1}{j+1/2} - C(Z, j, l) \right] \quad C \approx 0.6$$

1. Increases with nuclear charge Z.

2. Changes sign for higher angular momemtum j.

Methods of Atomic Calculations

N _{ve}	Relativistic Hartree-Fock +	Accuracy		
1	All-orders sum of dominating diagrams	0.1-1%		
2-6	Configuration Interaction + Many-Body Perturbation Theory	1-10%		
2-15	Configuration Interaction	10-20%		
These methods cover all periodic system of elements				

They were used for many important problems:

- Test of Standard Model using Parity Violation in Cs,TI,Pb,Bi
- Predicting spectrum of Fr (accuracy 0.1%), etc.

Results of calculations (in cm⁻¹)

Anchor lines

Negative shifters

Atom	ω ₀	q
Mg I	35051.217	86
Mg II	35760.848	211
Mg II	35669.298	120
Si II	55309.3365	520
Si II	65500.4492	50
ALII	59851.924	270
AI III	53916.540	464
AI III	53682.880	216
Ni II	58493.071	-20

Also, many transitions in Mn II, Ti II, Si IV, C II, C IV, N V, O I, Ca I, Ca II, Ge II, O II, Pb II,Co II,...

Different signs and magnitudes of q provides opportunity to study systematic errors!

Atom	ω ₀	q	
Ni II	57420.013	-1400	C
Ni II	57080.373	-700	C
Cr II	48632.055	-1110	
Cr II	48491.053	-1280	
Cr II	48398.862	-1360	
Fe II	62171.625	-1300	

Positive shifters

Atom	ω ₀	q	
Fe II	62065.528	1100	
Fe II	42658.2404	1210	
Fe II	42114.8329	1590	
Fe II	41968.0642	1460	
Fe II	38660.0494	1490	
Fe II	38458.9871	1330	
Zn II	49355.002	2490	
Zn II	48841.077	1584	

4.2 Astrophysical constraints:

Quasars - probing the universe back to much earlier times



Alkali Doublet Method

(Bahcall, Sargent; Varshalovich, Potekhin, Ivanchik, et al)

Fine structure interval

$$\Delta_{FS} = E(p_{3/2}) - E(p_{1/2}) = A(Z\alpha)^2$$

If Δ_z is observed at red shift z and Δ_0 is FS measured on Earth then

$$\frac{\Delta \alpha}{\alpha} = \frac{1}{2} \left(\frac{\Delta_{z}}{\Delta_{0}} - 1 \right)$$

Ivanchik *et al*, 1999: $\Delta \alpha / \alpha = -3.3(6.5)(8) \times 10^{-5}$. Murphy *et al*, 2001: $\Delta \alpha / \alpha = -0.5(1.3) \times 10^{-5}$.

Variation in the fine structure constant?: Recent results and the future

The alkali doublet (AD) method

- The AD method is simple ... but inefficient.
- The common S ground state in ADs has maximal relativistic corrections!



(a) Silv alkali doublet

Request for laboratory measurements: shopping list arxiv: physics/0408017

- More accurate measurements of UV transition frequencies
- Measurements of isotope shifts

Cosmological evolution of isotope abundances in the Universe:

a). Systematics for the variation of $\boldsymbol{\alpha}$

b). Test of theories of nuclear reactions in stars and supernovae

 Oscillator strengths to fit column densities Probing the variability of α with QSO absorption lines

A new method



Michael Murphy, UNSW

Variation of fine structure constant $\boldsymbol{\alpha}$

Many-Multiplet Method

Relativistic correction to electron energy E_n :

$$\Delta_n = \frac{E_n}{\nu} (Z\alpha)^2 \left[\frac{1}{j+1/2} - C(Z, j, l) \right] \quad C \approx 0.6$$

1. Increases with nuclear charge Z.

2. Changes sign for higher angular momemtum j.

Probing the variability of α with OSO absorption lines **Procedure:**

- 1. Compare heavy (Z~30) and light (Z<10) atoms, OR
- 2. Compare s ____ p and d ____ p transitions in heavy atoms.

Shifts can be of opposite sign.

Basic fo

ormula:
$$\mathbf{E}_{z} = \mathbf{E}_{z=0} + q \left[\left(\frac{\alpha_{z}}{\alpha_{0}} \right)^{z} - 1 \right]$$

 $E_{z=0}$ is the laboratory frequency. 2^{nd} term is non-zero only if α has changed. q is derived from atomic calculations. (Method: frequencies of different lines are computed for different values of α). Relativistic shift of the q = Q + K(L.S)K is the spin-orbit splitting central line in the multiplet parameter. Q~10K Numerical examples: (units = cm⁻¹) CrII MgII Fell Z=26 (s \rightarrow p) FeII 2383A: $\omega_0 = 38458.987(2) + 1449x$ Z=12 (s \rightarrow p) MgII 2796A: $\omega_0 = 35669.298(2) + 120x$ Z=24 (d \rightarrow p) CrII 2066A: $\omega_0 = 48398.666(2) - 1267x$ wavenumber $\mathbf{x} = (\alpha_z / \alpha_0)^2 - 1$ MgII "anchor"

F

Г

Many Multiplet Method

(Dzuba, Flambaum, Webb)



 α_2

Advantages:

Order of magnitude gain in sensitivity

- Statistical: all lines are suitable for analysis
- •Observe all unverse (up to z=4.2)
- Many opportunities to study systematic errors

Quasar absorption spectra





Quasar absorption spectra





One needs to know $E(\alpha^2)$ for each line to do the fitting

Use atomic calculations to find $\omega(\alpha)$.

For
$$\alpha$$
 close to $\alpha_0 \quad \omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$

q is found by varying α in computer codes:

$$q = d\omega/dx = [\omega(0.1) - \omega(-0.1)]/0.2, x = \alpha^2/\alpha_0^2 - 1$$

$\alpha = e^2/2 \varepsilon_0 hc = 0$ corresponds to nonrelativistic limit (infinite c).
Methods were used for many important problems:

- Test of Standard Model using Parity Violation in Cs,TI,Pb,Bi
- Predicting spectrum of Fr (accuracy 0.1%), etc.

Probing the variability of α with QSO absorption lines

To find dependence of atomic transition frequencies on α we have performed calculations of atomic transition frequencies for different values of α .

- 1. Zero Approximation Relativistic Hartree-Fock method: energies, wave functions, Green's functions
- 2. Many-body perturbation theory to calculate effective Hamiltonian for valence electrons including self-energy operator and screening; perturbation $\longrightarrow V = H - H_{HF}$



3. Diagonalization of the effective Hamiltonian

Probing the variability of α with QSO absorption lines

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- 1. Zero Approximation Relativistic Hartree-Fock method: energies, wave functions, Green's functions
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from core

3. Diagonalization of the effective Hamiltonian

Test: Energy levels in Mg II to 0.2% accuracy

Correlation potential method

[Dzuba,Flambaum,Sushkov (1989)]

- Zeroth-order: relativistic Hartree-Fock. Perturbation theory in difference between exact and Hartree-Fock Hamiltonians.
- Correlation corrections accounted for by inclusion of a "correlation potential" ∑:

$$V_{HF} \rightarrow V_{HF} + \Sigma$$

In the lowest order Σ is given by:



 External fields included using Time-Dependent Hartree-Fock (RPAE core polarization)+correlations

The correlation potential

Use the Feynman diagram technique to include three classes of diagrams to all orders:



The correlation potential

Use the Feynman diagram technique to include three classes of diagrams to all orders:





2. hole-particle interaction



3. nonlinear-in- Σ corrections



Atoms with several valence electrons: CI+MBPT

[Dziba, Flambailm, Kozlov (1996)]

CI Hamiltonian: $\Sigma_{i} h_{i} + \Sigma_{ij} e^{2} / t_{ij}$ $h = c \alpha p + (\beta - 1)mc^{2} - Ze^{2} / t + V_{core}$ CI+MBPT Hamiltonian: $h -> h + \Sigma_{j}; e^{2} / t_{ij} -> e^{2} / t_{ij} + \Sigma_{2}$

MBPT is used to calculate core-valence correlation operator X(r,r',E)



Atoms of interest

Z	Atom / Ion	Transitions	N _{ve} ¹
6	C I, C II, C III	p-s	4, 3, 2
8	01	p-s	4
11	Na I	s-p	1
12	Mg I, Mg II	s-p	2, 1
13	AI II, AI III	s-p	2, 1
14	Si II, Si IV	p-s	3, 1
16	S II	s-p	4
20	Ca II	s-p	1
22	Ti II	s-p, d-p	3
24	Cr II	d-p	5
25	Mn II	s-p, d-p	1
26	Fe II	s-p, d-p	7
28	Ni II	d-p	9
30	Zn II	s-p	1

 $^{1}N_{ve}$ – number of valence electrons

Methods of Atomic Calculations

N _{ve}	Relativistic Hartree-Fock +	Accuracy	
1	All-orders sum of dominating diagrams	0.1-1%	
2-6	Configuration Interaction + Many-Body Perturbation Theory	1-10%	
2-15	Configuration Interaction	10-20%	
These methods cover all periodic system of elements			

They were used for many important problems:

- Test of Standard Model using Parity Violation in Cs,TI,Pb,Bi
- Predicting spectrum of Fr (accuracy 0.1%), etc.

Relativistic shifts-doublets

Energies of "normal" fine structure doublets as functions of α^2





Relativistic shifts-triplets

Energies of "normal" fine structure triplets as functions of α^2





Fine structure anomalies and level crossing

Energies of strongly interacting states as functions of α^2





Implications to study of α variation

- Not every energy interval behaves like $\Delta E = A + B(Z\alpha)^2$.
- Strong enhancement is possible (good!).
- Level crossing may lead to instability of calculations (bad!).

Problem: level pseudo crossing

Energy levels of Ni II as functions of α^2



Values of *q=dE/dα*² are sensitive to the position of level crossing

Problem: level pseudo crossing

Energy levels of Ni II as functions of α^2



Values of $q=dE/d\alpha^2$ are sensitive to the position of level crossing

Solution: matching experimental gfactors

hyperfine= $\alpha^2 g_p m_e / M_p$ atomic units Rotation= m_e/M_p atomic units

Variation in the fine structure constant?: Recent results and the future

Radio constraints:

- > Hydrogen hyperfine transition at $\lambda_{H} = 21$ cm.
- Molecular rotational transitions CO, HCO+, HCN, HNC, CN, CS ...
- > $\omega_{\rm H}/\omega_{\rm M} \propto \alpha^2 g_{\rm P}$ where $g_{\rm P}$ is the proton magnetic *g*-factor.

$$g_{p} = g_{p} \left(\frac{m_{q}}{\Lambda_{qcD}} \right)$$

Probing the variability of α with QSO absorption lines **Procedure:**

- 1. Compare heavy (Z~30) and light (Z<10) atoms, OR
- 2. Compare s → p and d → p transitions in heavy atoms.

Shifts can be of opposite sign.

Basic formula:

a:
$$\mathbf{E}_{z} = \mathbf{E}_{z=0} + \mathbf{q} \left[\left(\frac{\alpha_{z}}{\alpha_{0}} \right)^{2} - 1 \right]$$

 $E_{z=0}$ is the laboratory frequency. 2^{nd} term is non-zero only if α has changed. q is derived from atomic calculations. (Method: frequencies of different lines are computed for different values of α).

Relativistic shift of the
central line in the multipletq = Q + K(L.S)
K is the spin-orbit splitting
parameter. $Q \sim 10K$ Z=26 (s \rightarrow p)FeII 2383A: $\omega_0 = 38458.987(2) + 1449x$
Z=12 (s \rightarrow p)MgII 2796A: $\omega_0 = 35669.298(2) + 120x$
Z=24 (d \rightarrow p)Crill 2066A: $\omega_0 = 48398.666(2) - 1267x$
x = $(\alpha_z/\alpha_0)^2 - 1$ MgII "matcher"

Michael Murphy, UNSW











Murphy et al, 2003: Keck telescope, 143 systems, 23 lines, 0.2<z<4.2
 Δα/α=-0.54(0.12) x 10⁻⁵

 Quast et al, 2004: VL telescope, 1 system, Fe II, 6 lines, 5 positive *q*-s, one negative *q*, *z*=1.15
 Δα/α=-0.4(1.9)(2.7) × 10⁻⁶

 Molaro et al 2007 -0.12(1.8) × 10⁻⁶, z=1.84 5.7(2.7)× 10⁻⁶

Srianand et al, 2004: VL telescope, 23 systems, 12 lines, Fe II, Mg I, Si II, Al II, 0.4<z<2.3
 Δα/α=-0.06(0.06) x 10⁻⁵

Murphy et al 2007 $\Delta \alpha / \alpha = -0.64(0.36) \times 10^{-5}$ Further revision may be necessary.

Probing the variability of a with QSO absorption lines Potential systematic effects:

> Laboratory wavelength errors: New, mutually consistent laboratory spectra from Imperial College, Lund University and NIST

> Data quality variations: Can only produce systematic shifts if combined with laboratory wavelength errors

> Heliocentric velocity variation: Smearing in velocity space is degenerate with fitted redshift parameters

> Isotopic ratio shifts: Very small effect possible if evolution of isotopic ratios allowed

> Hyperfine structure shifts: same as for isotopic shifts

> Magnetic fields: Large scale fields could introduce correlations in $\Delta \alpha/\alpha$ for neighbouring QSO site lines (if QSO light is polarised) - extremely unlikely and huge fields required

> Wavelength miscalibration: mis-identification of ThAr lines or poor polynomial fits could lead to systematic miscalibration of wavelength scale

Temperature changes during observations: Refractive index changes between ThAr and QSO exposures – random error

> Line blending: Are there ionic species in the clouds with transitions close to those we used to find $\Delta \alpha / \alpha$?

> Atmospheric refraction effects: Different angles through optics for blue and red light – can only produce positive $\Delta \alpha / \alpha$ at low redshift

> Instrumental profile variations: Intrinsic IP variations along spectral direction of CCD?

Possible systematic effect: i sotopic ratio evolution Different isotope abundancies -> shift of line. we calculated isotopic shifts for Hg I, si II (Pas), si IV, Zn II. However, calculations are too complicated for open d-shell atoms CrII, Fell, NiII, (also Sill sip-sp?) - in progress. Measure, please !!! "Conspiracy" of isotopic shifts and isotopic abundances ?. Line removal test.

Probing the variability of α with QSO absorption lines

Checks on general, unknown systematics:

- Line removal: In each system, remove each transition and iterate to find Δα/α again. Compare the Δα/α's before and after line removal. We have done this for all species and see no inconsistencies. Tests for: Lab wavelength errors, line blending, isotopic ratio and hyperfine structure variation.
- Positive-negative shifter test: Find the subset of systems that contain an anchor line, a positive shifter AND a negative shifter. Remove each type of line collectively and recalculate Δα/α. Results: subset contains 12 systems (only in high z sample) No lines removed: Δα/α = (-1.31 ± 0.39) × 10⁻⁵ Anchors removed: Δα/α = (-1.49 ± 0.44) × 10⁻⁵ +ve-shifters removed: Δα/α = (-1.54 ± 1.03) × 10⁻⁵ -ve-shifters removed: Δα/α = (-1.41 ± 0.65) × 10⁻⁵





Two sets of line pairs

- $1.\delta\alpha < 0$ imitated by compression of the spectrum
- 2. $\delta \alpha < 0$ imitated by expansion of the spectrum
- Both sets give $\delta \alpha < 0$!

New interpretation: Spatial variation

Northern+(new)Southern hemisphere data: Linear variation with distance along some direction, $\alpha(x)=\alpha(0)+kx$,

$$x=r \cos(\phi), r=ct (Gly),$$

 $\Delta \alpha / \alpha = 1.10(0.25) \ 10^{-6} \ r \cos(\phi)$

gradient direction 17.6(0.6) h, -58(6)°

 4.2σ deviation from zero. Data from two largest telescopes, Keck and VLT, give consistent results. 300 systems.

New interpretation: Spatial variation

Northern+(new)Southern hemisphere data: Linear variation with distance along some direction, $\alpha(x)=\alpha(0)+kx$, $x=r\cos(\phi)$, r=ct (Gly), $\Delta\alpha/\alpha = 1.2 \ 10^{-6} \ r\cos(\phi)$ dipole 4.1 σ deviation from zero. Data from two largest

telescopes, Keck and VLT, give consistent results. 300 systems.

Results for m_q/Λ_{QCD} and m_e/m_p Big Bang Nucleosynthsis data and H₂ molecule data are consitent with the direction of the dipole.

4.1 σ evidence for a $\Delta \alpha / \alpha$ dipole from VLT + Keck



Julian King, UNSW

The Keck & VLT dipoles point in the same direction



Julian King, UNSW

m_e / M_p limit from NH₃-2 systems

Inversion spectrum: exponentially small"quantum tunneling" frequency ω_{inv} =W exp(-S(m_e / M_p)) ω_{inv} is exponentially sensitive to m_e / M_p Laboratory measurements proposed (Veldhoven et al)

Flambaum,Kozlov PRL 2007 First enhanced effect in quasar spectra $\Delta(m_e / M_p) / (m_e / M_p)=-0.6(1.9)10^{-6}$ No variation z=0.68, 6.5 billion years ago, -1(3)10⁻¹⁶ /year

More accurate measurements Murphy, Flambaum, Henkel, Muller. Science 2008 -0.74(0.47)(0.76)10⁻⁶ Henkel et al AA 2009 z=0.87 <1.4 10⁻⁶ 3 σ

Levshakov, Molaro, Kozlov2008 our Galaxy 0.5(0.14)10-7

Hydrogen molecule - 4 systems

 $\Delta(m_e / M_p) / (m_e / M_p) =$ 3.3(1.5) 10⁻⁶ r cos(ϕ) gradient direction 16.7(1.5) h, $-62(5)^{\circ}$ consistent with α gradient direction 17.6(0.6) h, -58(6)° If we assume the same direction **2.6(1.3)** 10^{-6} r cos(ϕ) 4% by chance Big Bang nucleosynthesis: dependence on quark mass

- Flambaum, Shuryak 2002
- Flambaum, Shuryak 2003
- Dmitriev, Flambaum 2003
- Dmitriev, Flambaum, Webb 2004
- Coc, Nunes, Olive, Uzan, Vangioni 2007
- Dent, Stern, Wetterich 2007
- Flambaum, Wiringa 2007
- Berengut, Dmitriev, Flambaum 2009

Deuterium abundance – 7 points

Big Bang Nucleosynthsis data give direction of the gradient in the deuterium abundance consitent with the direction of the α gradient. However, the amplitude of the relative spatial variation 0.0045(35) is not statistically significant. This would result in relative variation of X=m_q/ Λ_{QCD}

∆X/X=0.0013(10) r cos(\u00fc)

 $\Delta \alpha / \alpha = 0.003(3) \operatorname{r} \operatorname{cos}(\phi)$

Compare with QSO

 $\Delta \alpha / \alpha = 1.10(0.25) \ 10^{-6} \ r \cos(\phi)$
Gradient α points down



Oklo natural nuclear reactor

n+¹⁴⁹Sm capture cross section is dominated by $E_r = 0.1 \text{ eV}$ resonance. Shlyakhter-limit on $\Delta \alpha / \alpha$ two billion years ago

Our QCD/nuclear calculations $\Delta E_r = 10 \text{ Mev} \Delta X_q / X_q - 1 \text{ MeV} \Delta \alpha / \alpha$ $X_q = m_q / \Lambda_{QCD}$, enhancement 10 MeV/0.1 eV=10⁸

Galaxy moves 552 km/s relative to CMB, $cos(\phi)=0.23$ Dipole in space: $\Delta E_r = (10 \text{ R} - 1) \text{ meV}$

Fujii et al $|\Delta E_r| < 20 \text{ MeV}$ Gould et al, $-12 < \Delta E_r < 26 \text{ meV}$ Petrov et al $-73 < \Delta E_r < 62 \text{ meV}$

Consequences for atomic clocks

- Sun moves 369 km/s relative to CMB cos(φ)=0.1 This gives average laboratory variation Δα/α =1.5 10⁻¹⁸ cos(φ) per year
- Earth moves 30 km/s relative to Sun-1.6 10 $^{-20}$ cos(ω t) annual modulation

Big Bang Nucleosynthesis: Dependence on m_q/Λ_{QCD}

- ²H 1+7.7x=1.07(15) x=0.009(19)
- ⁴He 1-0.95x=1.005(36) x=-0.005(38)
- ⁷Li 1-50x=0.33(11) x=0.013(02)

Final result

x= $\Delta X_q/X_q$ =0.013 (02), $X_q=m_q/\Lambda_{QCD}$ If we fit spatial dipole, the direction is the same as in quasar data. Measurements m_e / M_p or m_e / Λ_{QCD}

 Tsanavaris,Webb,Murphy,Flambaum, Curran PRL 2005
 Hyperfine H/optical , 9 quasar absorption systems with Mg,Ca,Mn,C,Si,Zn,Cr,Fe,Ni
 Measured X=α² g_p m_e / M_p
 ΔX/X=0.6(1.0)10⁻⁵ No variation

m_e / M_p limit from NH_3

Inversion spectrum: exponentially small"quantum tunneling" frequency ω_{inv} =W exp(-S) S=(m_e / M_p)^{-0.5} f(E_{vibration}/E_{atomic}) , E_{vibration}/E_{atomic} =const (m_e / M_p)^{-0.5} ω_{inv} is exponentially sensitive to m_e / M_p Flambaum,Kozlov PRL 2007 First enhanced effect in quasar spectra, 5 times Δ (m_e / M_p)/(m_e / M_p)=-0.6(1.9)10⁻⁶ No variation z=0.68, 6.5 billion years ago, -1(3)10⁻¹⁶ /year

More accurate measurements Murphy, Flambaum, Henkel, Muller. Science 2008 - $0.74(0.47)(0.76)10^{-6}$ Henkel et al AA 2009 z=0.87 <1.4 10⁻⁶ 3 σ

Levshakov, Molaro, Kozlov2008 our Galaxy 0.5(0.14)10-7



Measurements m_e / M_p or m_e / Λ_{QCD}

Reinhold, Buning, Hollenstein, Ivanchik,
 Petitjean, Ubachs PRL 2006, H₂ molecule, 2 systems

 $\Delta(m_e$ / M_p)/ (m_e / $M_p)$ =-2.4(0.6)10^-5 $\,$ Variation 4 σ ! Higher redshift, z=2.8

Space-time variation? Grand Unification model? 2008 Wendt,Reimers <4.9 10⁻⁵ 2008 Webb et al 0.26(0.30)10⁻⁵

Oklo natural nuclear reactor

 $n+^{149}$ Sm capture cross section is dominated by $E_r = 0.1 \text{ eV}$ resonance Shlyakhter;Damour,Dyson;Fujii et al Limits on variation of alpha

Flambaum,Shuryak 2002,2003 Dmitriev,Flambaum 2003 Flambaum,Wiringa 2008 $\Delta E_r = 10 \text{ Mev} \Delta X_q / X_q - 1 \text{ MeV} \Delta \alpha / \alpha$ $X_q = m_q / \Lambda_{QCD}$, enhancement 10 MeV/0.1 eV=10⁸

2006 Gould et al, Petrov et al $|\Delta E_r| < 0.1 \text{eV}$, $|\Delta X/X| < 10^{-8}$ two billion years ago, 10^{-17} /year

There are non-zero solutions

Oklo natural nuclear reactor

1.8 billion years ago

 $n+^{149}Sm$ capture cross section is dominated by $E_r = 0.1 \text{ eV}$ resonance

Shlyakhter;Damour,Dyson;Fujii et al

 $\Delta E_r = 1 \text{ MeV } \Delta \alpha / \alpha$

Limits on variation of alpha

Oklo: limits on $X_q = m_q / \Lambda_{QCD}$

Flambaum, Shuryak 2002, 2003 Dmitriev, Flambaum 2003 Flambaum, Wiringa 2008 150 Sm $\Delta E_r = 10 \text{ MeV} \Delta X_a/X_a - 1 \text{ MeV} \Delta \alpha/\alpha$

Limits on $x=\Delta X_q/X_q - 0.1 \Delta \alpha / \alpha$ from Fujii et al $|\Delta E_r| < 0.02 \text{ eV} |x| < 2.10^{-9}$ Petrov et al $|\Delta E_r| < 0.07 \text{ eV} |x| < 8.10^{-9}$ Gould et al $|\Delta E_r| < 0.026 \text{ eV} |x| < 3.10^{-9}$, <1.610⁻¹⁸ y⁻¹ There is second, non-zero solution x=1.0(1) 10⁻⁸

Atomic clocks

Cesium primary frequency standard:

HFS of 6s:
$$F=4 - v = 9 \ 192 \ 631 \ 770 \ Hz F=3 - v = 9 \ 192 \ 631 \ 770 \ Hz$$

Also: Rb, Cd⁺, Ba⁺, Yb⁺, Hg⁺, etc.

E.g. v(Hg⁺) = 40 507 347 996.841 59(14)(41) Hz (D. J. Berkeland *et al*, 1998).

Optical frequency standards:

Ζ	Atom	Transition	Frequency	Source
20	Са	¹ S ₀ - ³ P ₁	455 986 240 494 144(5.3) Hz	Degenhardt et al, 2005
38	Sr⁺	¹ S ₀ - ³ P ₁	434 829 121 311(10) kHz	Ferrari et al, 2003
49	ln⁺	¹ S ₀ - ³ P ₀	1 267 402 452 899 920(230) Hz	von Zanthier et al, 2005
70	Yb⁺	² S _{1/2} - ² F _{7/2}	642 121 496 772 300(600) Hz	Hosaka et al, 2005

Also: H, Al⁺, Sr, Ba⁺, Yb, Hg, Hg⁺, Tl⁺, Ra⁺, etc.

Accuracy about 10⁻¹⁵ can be further improved to 10⁻¹⁸!

Atomic clocks:

Comparing rates of different clocks over long period of time can be used to study time variation of fundamental constants!

Optical transitions: α

Microwave transitions: α , $(m_e, m_q)/\Lambda_{QCD}$

Advantages:

- Very narrow lines, high accuracy of measurements.
- Flexibility to choose lines with larger sensitivity to variation of fundamental constants.

• Simple interpretation (local time variation).

Calculations to link change of frequency to change of fundamental constants:

Optical transitions: <u>atomic calculations</u> (as for quasar absorption spectra) for many narrow lines in Al II, Ca I, Sr I, Sr II, In II, Ba II, Dy I, Yb I, Yb II, Yb III, Hg I, Hg II, TI II, Ra II, ThIV $\omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$

Microwave transitions: hyperfine frequency is sensitive to nuclear magnetic moments and nuclear radii

We performed atomic, nuclear and QCD calculations of powers κ , β for H,D,Rb,Cd⁺,Cs,Yb⁺,Hg⁺ V=C(Ry)(m_e/M_p) $\alpha^{2+\kappa}$ (m_q/ Λ_{QCD}) $^{\beta}$, $\Delta\omega/\omega=\Delta V/V$ **Calculations** to link change of frequency to change of fundamental constants:

Optical transitions: <u>atomic calculations</u> (as for quasar absorption spectra) for many narrow lines in Al II, Ca I, Sr I, Sr II, In II, Ba II, Dy I, Yb I, Yb II, Yb III, Hg I, Hg II, TI II, Ra II ... $\omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$

Microwave transitions: hyperfine frequency is sensitive to α , nuclear magnetic moments and nuclear radii

We performed atomic, nuclear and QCD calculations

of powers κ , β for H,D,He,Rb,Cd⁺,Cs,Yb⁺,Hg⁺... $V=C(Ry)(m_e/M_p)\alpha^{2+\kappa} (m_a/\Lambda_{QCD})^{\beta}, \Delta\omega/\omega=\Delta V/V$ ¹³³Cs: $\kappa = 0.83$, $\beta = 0.002$ Cs standard is insensitive to variation of m_{d}/Λ_{OCD} ! ⁸⁷Rb: $\kappa = 0.34$, $\beta = -0.02$ ¹⁷¹Yb+: $\kappa = 1.5, \beta = -0.10$ ¹⁹⁹Hg+: $\kappa = 2.28$, $\beta = -0.11$ ¹H: $\kappa = 0, \beta = -0.10$

Complete Table in Phys.Rev.A79,054102(2009)

Results for variation of fundamental constants

Source	Clock ₁ /Clock ₂	$d\alpha/dt/\alpha(10^{-16} {\rm yr}^{-1})$	
Blatt <i>et al</i> , 2007	Sr(opt)/Cs(hfs)	-3.1(3.0)	
Fortier et al 2007	Hg+(opt)/Cs(hfs)	-0.6(0.7) ^a	
Rosenband et al08	Hg+(opt)/Al+(opt)	-0.16(0.23)	
Peik <i>et al</i> , 2006	Yb+(opt)/Cs(hfs)	4(7)	
Bize <i>et al</i> , 2005	Rb(hfs)/Cs(hfs)	1(10) ^a	

^aassuming $m_{q,e}/\Lambda_{QCD}$ = Const

Combined results: $d/dt \ln \alpha = -1.6(2.3) \times 10^{-17} \text{ yr}^{-1}$ $d/dt \ln(m_q/\Lambda_{QCD}) = 3(25) \times 10^{-15} \text{ yr}^{-1}$ $m_e /M_p \text{ or } m_e/\Lambda_{QCD} -1.9(4.0) \times 10^{-16} \text{ yr}^{-1}$



Transition frequency shifts with fine-structure constant variation for Yb II

S. G. Porsev,^{1,2} V. V. Flambaum,¹ and J. R. Torgerson³ ¹School of Physics, University of New South Wales, Sydney, New South Wales 2052, Australia ²Petersburg Nuclear Physics Institute, Gatchina, Leningrad District 188300, Russia ³Los Alamos National Laboratory, University of California, Physics Division, P.O. Box 1663, Los Alamos, New Mexico 87545, USA (Received 20 July 2009; published 7 October 2009)

Larger q in Yb II

- Transition from ground state f^{14} 6s ${}^2S_{1/2}$ to metastable state f^{13} 6s ${}^2{}^2F_{7/2}$ q₁=-60 000
- For transitions from metastable state $f^{13}6s^2 {}^2F_{7/2}$ to higher metastable states q_2 are positive and large, up to 85 000 Difference $q=q_2 - q_1$ may exceed 140 000,
- so the sensitivity to alpha variation using comparison of two transitions in Yb II exceeds that in HgII/All comparison (measurements at NIST) 2.7 times.
- Shift of frequency difference is 2.7 times larger

Porsev, Flambaum, Torgerson

Largest q in multiply charged ions, narrow lines

q increases as $Z^2 (Z_i+1)^2$

To keep frequencies in optical range we use configuration crossing as a function of Z

Crossing of 5f and 7s Th IV: q_1 =-75 300

Crossing of 4f and 5s Sm15+, Pm14+, Nd 13+ Difference $q=q_2 - q_1$ is 260 000 5 times larger than in Hg II/Al II Relative sensitivity enhancement up to 500 Berengut, Dzuba, Flambaum, Porsev



FIG. 2. Dirac-Fock ionisation energies of 5s (solid) and $4f_{7/2}$ (dashed) levels for the Ag isoelectronic sequence.

arXiv:1007.1068

Enhancement of relative effect

Dy: $4f^{10}5d6s = 19797.96... \text{ cm}^{-1}$, $q = 6000 \text{ cm}^{-1}$ $4f^{9}5d^{2}6s = 19797.96... \text{ cm}^{-1}$, $q = -23000 \text{ cm}^{-1}$ Interval $\Delta \omega = 10^{-4} \text{ cm}^{-1}$

Relative enhancement $\Delta\omega/\omega_0 = 10^8 \Delta\alpha/\alpha$

Measurement Berkeley $d\ln\alpha/dt = -2.9(2.6) \times 10^{-15} \text{ yr}^{-1}$

Close narrow levels in molecules and nucleus ²²⁹Th

Dysprosium miracle

Dy: $4f^{10}5d6s = 19797.96... \text{ cm}^{-1}$, $q = 6000 \text{ cm}^{-1}$ $4f^{9}5d^{2}6s = 19797.96... \text{ cm}^{-1}$, $q = -23000 \text{ cm}^{-1}$ Interval $\Delta \omega = 10^{-4} \text{ cm}^{-1}$

Our calculations: Enhancement factor $K = 10^8$ (!), i.e. $\Delta \omega / \omega_0 = 10^8 \Delta \alpha / \alpha$

Measurements (Berkeley,Los Alamos) $d \ln \alpha / dt = -2.7(2.6) \times 10^{-15} \text{ yr}^{-1}$

Problem: states are not narrow! There are close narrow levels in molecules.

More suggestions ...

Atom	State ₁		State ₂		K
Ce I	⁵ H ₃	2369.068	¹ D ₂	2378.827	2000
	³ H ₄	4762.718	³ D ₂	4766.323	13000
Nd I	⁵ K ₆	8411.900	7L ₅	8475.355	950
Nd I	⁷ L ₅	11108.813	⁷ K ₆	11109.167	10 ⁵
Sm I	⁵ D ₁	15914.55	⁷ G ₂	12087.17	300
Gd II	⁸ D _{11/2}	4841. 106	¹⁰ F _{9/2}	4852.304	1800
Tb I	⁶ H _{13/2}	2771.675	⁸ G _{9/2}	2840.170	600

Enhancement in molecular clocks

- DeMille et al 2004, 2008 enhancement in Cs_2 , cancellation between electron excitation and vibration energies
- Flambaum 2006 Cancellations between rotational and hyperfine intervals

 $\Delta \omega / \omega_0 = K \Delta \alpha / \alpha$ Enhancement K = 10² - 10³

Flambaum, Kozlov 2007 Cancellations between fine structure and vibrations

 $\Delta \omega / \omega_0 = K (\Delta \alpha / \alpha - 1/4 \Delta \mu / \mu)$ Enhancement K = 10⁴ - 10⁵

Enhancement in molecular clocks

- DeMille 2004, DeMille et al 2008 enhancement in Cs₂, cancellation between electron excitation and vibration energies
- Flambaum 2006 Cancellations between rotational and hyperfine intervals in very narrow microwave transitions in LaS, LaO, LuS,LuO, YbF, etc.

 $ω_0 = E_{rotational} - E_{hyperfine} = E_{hyperfine} / 100 - 1000$ $Δω/ω_0 = K Δα/α$ Enhancement K = 10² - 10³

Cancellation between fine structure and vibrations in molecules

Flambaum, Kozlov PRL2007 K = 10⁴ - 10⁵,

SiBr, Cl_2^+ ... microwave transitions between narrow excited states, sensitive to α and $\mu=m_e/M_p$

$$\omega_0 = E_{\text{fine}} - E_{\text{vibrational}} = E_{\text{fine}} / K$$

 $\Delta \omega / \omega_0 = K (\Delta \alpha / \alpha - 1/4 \Delta \mu / \mu)$
Enhancement $K = 10^4 - 10^5$

E _{fine} is proportional to $Z^2\alpha^2$

 $E_{vibrational} = n\omega$ is proportional to $n\mu^{0.5}$, n=1,2,...

Enhancement for all molecules along the lines $Z(\mu,n)$

Shift 0.003 Hz for $\Delta \alpha / \alpha = 10^{-16}$; width 0.01 Hz

Compare with Cs/Rb hyperfine shift 10⁻⁶ Hz

HfF⁺ K = 10^3 shift 0.1 Hz

Cancellation between fine structure and rotation in light molecules Bethlem, Bunning, Meijer, Ubach 2009 OH,OD,CN,CO,CH,LiH,... E_{fine} is proportional to $Z^2 \alpha^2$ $E_{rotational}$ is proportional to Lµ, L=0,1,2,... $\mu = m_e / M_p$ Enhancement for all molecules along the lines $Z(\mu,L)$

Nuclear clocks (suggested by Peik,Tamm 2003)

Very narrow UV transition between first excited and ground state in ²²⁹ Th nucleus Energy 7.6(5) eV, width 10^{-4} Hz Flambaum PRL2006 Nuclear/QCD estimate: Enhancement 10⁵, $\Delta \omega / \omega_0 = 10^5 (0.1 \Delta \alpha / \alpha + \Delta X_0 / X_0)$ $X_a = m_a / \Lambda_{QCD}$, Shift 10⁵ Hz for $\Lambda \alpha / \alpha = 10^{-15}$ Compare with atomic clock shift 1 Hz

 235 U energy 76 eV, width 6 10 $^{-4}$ Hz

Nuclear clocks

Peik, Tamm 2003: UV transition between first excited and ground state in ²²⁹Th nucleus Energy 7.6(5) eV, width 10⁻³ Hz. Perfect clock!

Flambaum 2006: Nuclear/QCD estimate- Enhancement **10**⁵

He,Re; Flambaum,Wiringa; Flambaum,Auerbach,Dmitriev; Hayes,Friar,Moller;Litvinova,Felmeier,Dobaczewski,Flambaum; $\Delta \omega = 10^{19}$ Hz ($\Delta \alpha / \alpha + 10 \Delta X_q / X_q$), $X_q = m_q / \Lambda_{QCD}$, Shift 10-100 Hz for $\Delta \alpha / \alpha = 10^{-18}$ Compare with atomic clock shift 0.001 Hz

Berengut, Dzuba, Flambaum, Porsev: Sensitivity to $\Delta \alpha / \alpha$ is expressed via isomeric shifts of ²²⁹Th atomic lines, frequency in ²²⁹Th - frequency in ²²⁹Th *. Measure, please!

Enhancement of relative effect

Dy: $4f^{10}5d6s = 19797.96... \text{ cm}^{-1}$, $q = 6000 \text{ cm}^{-1}$ $4f^{9}5d^{2}6s = 19797.96... \text{ cm}^{-1}$, $q = -23000 \text{ cm}^{-1}$ Interval $\Delta \omega = 10^{-4} \text{ cm}^{-1}$

Relative enhancement $\Delta \omega / \omega_0 = 10^8 \Delta \alpha / \alpha$

Measurement Berkeley $d\ln\alpha/dt = -2.9(2.6) \times 10^{-15} \text{ yr}^{-1}$

Close narrow levels in molecules

Conclusions

 Spatial gradient from quasar data provides alpha variation for atomic clocks due to Earth motion at the level 10⁻¹⁸ per year and 1 meV shift in Oklo resonance. One-two orders of magnitude improvement in the measurement accuracy is needed. Three orders for meteorites.

New systems with higher absolute sensitivity include:

- transitions between ground and metastable states in highly charged ions.
 Frequencies are kept in laser spectroscopy range due to the configuration crossing phenomenon. An order of magnitude gain.
- ²²⁹Th nucleus highest absolute enhancement (10⁵ times larger shift), UV transition 7eV.
- Many systems with relative enhancement due to transition between close levels: Dy atom, a number of molecules with narrow close levels,...
- Very weak indications for the spatial variation in H₂ quasar spectra and BBN abundance of deuterium. The same direction of the gradient!
- Search for anisotropy in CMB, expansion of the Universe, structure formation

Conclusions

 Spatial dipole in quasar data provides alpha variation for atomic clocks due to Earth motion at the level 10⁻¹⁸ per year.

New systems with higher absolute sensitivity include:

- transitions between metastable states in Yb II
- transitions between ground state and metastable state in Th 3+ and many highly charged ions. Frequencies are kept in laser spectroscopy range due to the configuration crossing phenomenon. An order of magnitude gain.
- ²²⁹Th nucleus highest absolute enhancement (10⁵ times larger shift), UV transition 7eV.
- Many systems with relative enhancement due to transition between close levels: Dy atom, a number of molecules with narrow close levels,...

Nuclear clocks

Peik, Tamm 2003: UV transition between first excited and ground state in ²²⁹Th nucleus. Energy 7.6(5) eV, width 10⁻⁴ Hz. Perfect clock!

Our nuclear/QCD calculations - Enhancement 10⁵

$$\begin{split} &\Delta\omega/\omega_0 = \ 10^5 \ (\ 0.1 \Delta\alpha/\alpha + \Delta X_q/X_q \) \\ &X_q = m_q/ \ \Lambda_{QCD} \ , \\ &\text{Shift 2000 Hz for } \Delta\alpha/\alpha = 10^{-16} \\ &\text{Compare with atomic clock shift 0.1 Hz} \end{split}$$

Problem – to find this narrow transition using laser Search: Peik et al, Lu et al, Habs et al, DeMille et al, Beck et al

²²⁹Th: why enhancement?

 $\omega = \mathbf{Q} + \mathbf{E}_{pk} + \mathbf{E}_{so} = 7.6 \text{ eV}$ huge cancellations! **Q=Coulomb=100 KeV** 10⁻⁴ total Coulomb E_{so} =<V_s L S>=spin-orbit=-1.0 MeV **E**_{pk} =potential+kinetic=1 MeV **Extrapolation from light nuclei** $\Delta E_{pk} / E_{pk} = -1.4 \Delta m_a / m_a$ $\Delta E_{so} = -0.24 \Delta m_a / m_a$ $\Delta \omega / \omega_0 = 10^5 (0.14 \Delta \alpha / \alpha + 1.6 \Delta X_0 / X_0)$
Dependence on α

$\Delta \omega = Q \Delta \alpha / \alpha$

- Total Coulomb energy 10³ MeV in ²²⁹ Th
- Difference of moments of inertia between ground and excited states is 4%
- If difference in the Coulomb energy would be 0.01%, Q=100 KeV, estimate for the enhancement factor

 $Q/\omega_0 = 10^5 \text{ eV} / 7 \text{ eV} = 1.4 \ 10^4$

Enhancement in ²²⁹Th

 α X_a=m_a/ Λ_{QCD} Flambaum 2006 ~10⁵ 0.5 10⁵ estimate Hayes, Frier 2007 0 impossible arguments He,Ren 2007 0.04 10⁵ 0.8 10⁵ rel.mean field Main effect (dependence of deformation on α) missed, change of mean-field potential only Dobaczewski et al 2007 0.15 10⁵ Hartree-Fock

preliminary

²²⁹Th: Flambaum, Wiringa 2007

 $\omega = E_{pk} + E_{so} = 7.6 \text{ eV}$ huge cancellations! E_{so} =<V_s L S>=spin-orbit=-1.04 MeV **E**_{pk} =potential+kinetic=1 MeV **Extrapolation from light nuclei** $\Delta E_{pk}/E_{pk}$ =-1.4 $\Delta m_a/m_a$ $\Delta E_{so}/E_{so} = -0.24 \Delta m_a/m_a$ $\Delta \omega / \omega_0 = 1.6 \ \mathbf{10^5} \ \Delta X_0 / X_0$

Difference of Coulomb energies

 $\Delta \omega = Q \Delta \alpha / \alpha$ Hayes, Frier, Moller <30 Kev He,Ren 30 KeV Flambaum, Auerbach, Dmitriev -500 Kev < Q < 1500 KeV Litvinova, Feldmeier, Dobaczewski. Flambaum -300 Kev < Q < 450 KeV

Sensitivity to $\Delta \alpha$ may be obtained from measurements

$\Delta \omega = Q \Delta \alpha / \alpha$

Berengut, Dzuba, Flambaum, Porsev PRL 2009

Q/Mev=-506 $\Delta < r^2 > / < r^2 > + 23 \Delta Q_2 / Q_2$

Diffrence of squared charge radii Δ <r²> may be extracted from isomeric shifts of electronic transitions in Th atom or ions

Diffrence of electric quadrupole moments ΔQ_2 from hyperfine structure

Experimental progress in ²²⁹Th

- Transition energy measured in Livermore
 7.6 (5) eV instead of 3.5(1.0) eV
- Intensive search for direct radiation
 - Argonne
 - Peik et al,
 - Habs et al, ...

Ultracold atomic and molecular collisions. Cheng Chin, Flambaum PRL2006

Enhancement near Feshbach resonance. Variation of scattering length $\Delta a/a=K \Delta \mu/\mu$, K=10² – 10¹² $\mu=m_e/M_p$ Hart,Xu,Legere,Gibble Nature 2007 Accuracy in scattering length 10⁻⁶ Evolution fundamental constants and their dependence on scalar and gravitational potential Fundamental constants depend on scalar field ϕ – dark energy, Higgs, dilaton, distance between branes, size of extra dimensions.

- Cosmological evolution of ϕ in space and time is linked to evolution of matter.
- Changes of Universe equation of state:
- Radiation domination, cold matter domination, dark energy domination-

Change of ϕ – change of $\alpha(\phi)$





Scalar charge-source of $\boldsymbol{\phi}$

- Massive bodies have scalar charge S proportional to the number of particles Scalar field ϕ =S/r, proportional to
 - gravitational potential GM/r -
- Variation of α proportional to gravitational potential

 $\delta \alpha / \alpha = K_{\alpha} \delta(GM/rc^2)$

Neutron star, white/brown dwarfs, galaxy, Earth, Sun – compare spectra, $\omega(\alpha)$

Dependence of fundamental constants on gravitational or scalar potential

Projects –atomic clocks at satellites in space or close to Sun (JPL project)

Earth orbit is elliptic,3% change in distance to Sun Fortier et al – Hg^{+(opt)}/Cs , Ashby et al -H/Cs

Flambaum, Shuryak : limits on dependence of α ,

 m_e / Λ_{QCD} and m_q / Λ_{QCD} on gravity

 $\delta \alpha / \alpha = K_{\alpha} \delta(GM/rc^2)$

 K_{α} +0.17 K_{e} =-3.5(6.0) 10⁻⁷

 K_{α} +0.13 K_{α} =2(17) 10⁻⁷

New results from Dy, Sr/Cs

Dysprosium $\delta \alpha / \alpha = K_{\alpha} \delta (GM/rc^2)$

Dy: $4f^{10}5d6s = 19797.96... \text{ cm}^{-1}$, $q = 6000 \text{ cm}^{-1}$ $4f^{9}5d^{2}6s = 19797.96... \text{ cm}^{-1}$, $q = -23000 \text{ cm}^{-1}$ Interval $\Delta \omega = 10^{-4} \text{ cm}^{-1}$

Enhancement factor K = 10⁸ , i.e. $\Delta\omega/\omega_0 = 10^8 \Delta\alpha/\alpha$

Measurements Ferrel et al 2007 K_{α} =-8.7(6.6) 10⁻⁶ K_{e} =4.9(3.9) 10⁻⁶ K_{q} =6.6(5.2) 10⁻⁶

Sr(optical)/Cs comparison : S.Blatt et al 2008

New best limits

 K_{α} =2.5(3.1) 10⁻⁶ K_{e} =-1.1(1.7) 10⁻⁶ K_{q} =-1.9(2.7) 10⁻⁶

Microwave clocks in optical lattice

Sr,Hg ,... in optical lattice. Optical clocks.
 Magic wavelength-cancellation of dynamical Stark shifts, very accurate optical frequencies.
 Katory, Kimble, Ye,...

- Hyperfine transitions, linear polarization no magic wavelength in atoms with valence s-electron: Cs , Rb,...
 There is magic wavelenght for atoms with p_{1/2} electrondue to hyperfine mixing p_{1/2}-p_{3/2} Al, Ga,...
 Beloy,Derevinako,Dzuba, Flambaum PRL 2009
- Circular polarisation- all wavelengths are magic for a certain direction of magnetic field "magic angle"
 Cs (primary standard), Rb,... PRL 2008

Conclusions

- Quasar data: MM method provided sensitivity increase 100 times. Anchors, positive and negative shifters-control of systematics. Keckvariation of α , VLT-?. Systematics or spatial variation.
- m_e/M_p : hyperfine H/optical, NH₃– no variation, H₂ variation 4 σ ? Space-time variation? Grand Unification model?
- Big Bang Nucleosynthesis: may be interpreted as a variation of $m_{\rm q}/\,\Lambda_{\rm QCD}$
- Oklo: sensitive to m_q / Λ_{QCD} , effect <10⁻⁸
- Atomic clocks: present time variation of α , m/ Λ_{QCD}
- Transitions between narrow close levels in atoms and molecules huge enhancement of the relative effect
- ²²⁹Th nucleus absolute enhancement (10⁵ times larger shift)
- Dependence of fundamental constants on gravitational potential

No variation for small red shift, hints for variation at high red shift

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