

The Liège Oscillation Code

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The Liège Oscillation Code

1. Inputs

Inputs

- global M and R
- radius: r
- mass: $m(r)$
- density: $\rho(r)$
- pressure: $P(r)$
- adiabatic: $\Gamma_1(r)$



dimensionless variables:

- $x = r/R$,
- q/x^3 (with $q=m/M$)
- $RP/GM\rho$,
- $4\pi R^3 \rho/M$
- Γ_1

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2. Non-radial adiabatic oscillations

The perturbation is described by 4 functions $Y(x)$ $Z(x)$ $U(x)$ $V(x)$

$$Y(x) = x^{1-\ell} \frac{\delta r}{R}$$

$$Z(x) = x^{-\ell} \frac{\delta P}{P}$$

$$U(x) = x^{-\ell} \frac{R \Phi'}{GM}$$

$$V(x) = x^{1-\ell} \left(\frac{R^2}{GM} \frac{d\Phi'}{dr} + \frac{4\pi\rho R^3}{M} \frac{\delta r}{R} \right).$$

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2. Non-radial adiabatic oscillations

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$$Y(x) = x^{1-\ell} \frac{\delta r}{R}$$

Lagrangian perturbation of the pressure

$$Z(x) = x^{-\ell} \frac{\delta P}{P}$$

$$U(x) = x^{-\ell} \frac{R \Phi'}{GM}$$

Eulerian perturbation of the
Gravitational potential

$$V(x) = x^{1-\ell} \left(\frac{R^2}{GM} \frac{d\Phi'}{dr} + \frac{4\pi\rho R^3}{M} \frac{\delta r}{R} \right).$$

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2. Non-radial adiabatic oscillations

A nonradial mode :

- $\vec{\delta r} = \sqrt{4\pi} \Re \left\{ \left[a(r) Y_{\ell m}(\theta, \phi) \vec{e}_r + b(r) \left(\frac{\partial Y_{\ell m}}{\partial \theta} \vec{e}_\theta + \frac{1}{\sin \theta} \frac{\partial Y_{\ell m}}{\partial \phi} \vec{e}_\phi \right) \right] e^{-i\sigma t} \right\}$
$$a(r)/R = x^{\ell-1} Y(x) \quad b(r)/R = \frac{x^{\ell-1}}{\omega^2} \left[U(x) + \frac{RP}{GM\rho} Z(x) + \frac{q}{x^3} Y(x) \right]$$
- $\frac{\delta P}{P} = \sqrt{4\pi} \Re \{ x^\ell Z(x) Y_{\ell m}(\theta, \phi) e^{-i\sigma t} \}$
- $\frac{R\Phi'}{GM} = \sqrt{4\pi} \Re \{ x^\ell U(x) Y_{\ell m}(\theta, \phi) e^{-i\sigma t} \}$
- $\frac{R^2}{GM} \frac{\partial \Phi'}{\partial r} = \sqrt{4\pi} \Re \left\{ x^{\ell-1} \left[V(x) - \frac{4\pi R^3 \rho}{M} Y(x) \right] Y_{\ell m}(\theta, \phi) e^{-i\sigma t} \right\}$

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2. Non-radial adiabatic oscillations

Differential equations

- $\frac{dY}{dx} = \frac{\ell+1}{x} \left\{ -Y + \frac{\ell}{\omega^2} \left(\frac{q}{x^3} Y + \frac{RP}{GM\rho} Z + U \right) \right\} - \frac{x}{\Gamma_1} Z$
- $\frac{dZ}{dx} = \frac{GM\rho}{RP} \left\{ \left(\omega^2 + 4\frac{q}{x^3} \right) \frac{Y}{x} + x \frac{q}{x^3} Z - \frac{V}{x} - \frac{\ell(\ell+1)}{x\omega^2} \frac{q}{x^3} \left(\frac{q}{x^3} Y + \frac{RP}{GM\rho} Z + U \right) \right\} - \frac{\ell}{x} Z$
- $\frac{dU}{dx} = \frac{1}{x} \left(V - \frac{4\pi R^3 \rho}{M} Y - \ell U \right)$
- $\frac{dV}{dx} = \frac{\ell+1}{x} (\ell U - V) + \frac{\ell(\ell+1)}{x\omega^2} \frac{4\pi R^3 \rho}{M} \left(\frac{q}{x^3} Y + \frac{RP}{GM\rho} Z + U \right)$

2. Non-radial adiabatic oscillations

- Boundary conditions at the centre: regularity

$$Y = \frac{\ell}{\omega^2} \left[\frac{q}{r^3} Y + \frac{RP}{GM\rho} Z + U \right]$$

$$V = \frac{4\pi R^3 \rho}{M} Y + \ell U$$

- Boundary conditions at the surface

- Gravitation potential $V + (\ell+1)U = 0$

- Pressure $\frac{\delta P}{P} + (4 + \omega^2) \frac{dr}{r} = 0$: Standard option in LOSC

$$\delta P = 0$$

3. Numerical techniques

3.1 Determination of mesh:

Option: Grid as based on eigenfunction asymptotic behaviour

3.2 Difference equations: 4th order scheme

We do not improve the solution with the Richardson extrapolation method

3.3 Inverse Iteration method (Keeley 1977, ApJ211,926)

To solve the eigenvalue problem $(A - \lambda B)\vec{y} = 0$ with $\lambda = \omega^2$

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3.2 Difference equations

4th order scheme based on the identity:

$$\vec{y}_i + \frac{h}{2}\vec{y}'_i + \frac{h^2}{12}\vec{y}''_i = \vec{y}_{i+1} - \frac{h}{2}\vec{y}'_{i+1} + \frac{h^2}{12}\vec{y}'''_{i+1} + O(h^5)$$

with $h = x_{i+1} - x_i$

If \vec{y} satisfies $\frac{d\vec{y}}{dx} = A(x)\vec{y}$

$$\left. \begin{aligned} & \left\{ 1 + \frac{h}{2}\alpha_i + \frac{h^2}{12}\beta_i \right\} \vec{y}_i = \left\{ 1 - \frac{h}{2}\alpha_{i+1} + \frac{h^2}{12}\beta_{i+1} \right\} \vec{y}_{i+1} \\ & \alpha = A \\ & \beta = A^2 + \frac{dA}{dx} \end{aligned} \right\}$$

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4. Outputs

- angular (σ) and dimensionless angular (ω) frequency
- mode order (k):
 - Phase diagramme : Scuflaire (1974, A&A 36, 107)
 - Lee (1985, PASJ 37,279) for more condensed stars
- $k = 0, 1, 2, \dots$ for p-modes with $l > 1$
- $k = 1, 2, 3, \dots$ for p-modes with $l = 0, 1$
- $k = -1, -2, -3, \dots$ for g-modes
- mode parity

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4. Outputs

- $f_T = E_{kin,V}/E_{kin}$
- $\langle x \rangle$
- $\Delta = \sqrt{\langle x^2 \rangle - \langle x \rangle^2}$
- rotational splitting coefficient $\sigma_{k\ell m} = \sigma_{k\ell}^0 + m \beta_{k\ell} \Omega$
- kernels $\sigma_{k\ell m} = \sigma_{k\ell}^0 + m \int K_{k\ell}(x) \Omega(x) dx$
- Eigenfunctions: $Y(x)$ $Z(x)$ $U(x)$ $V(x)$
and $Y'(x)$ $Z'(x)$ $U'(x)$ $V'(x)$

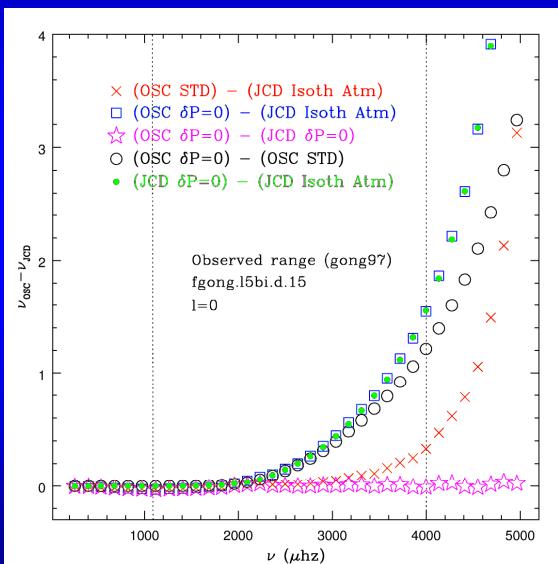
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5. Comparison with ADIPLS



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