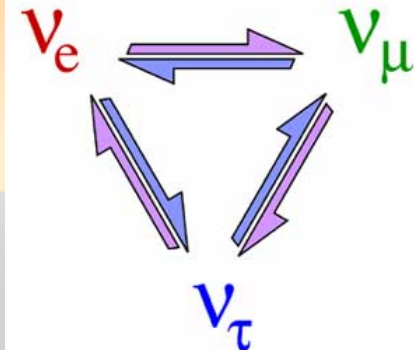
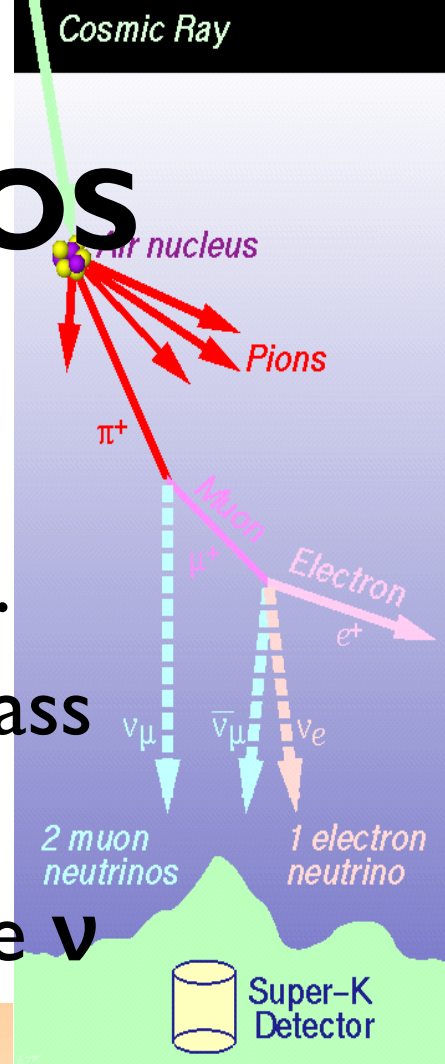


Merging Rates of the First Objects and Formation of the First Mini-Filaments in Models with Massive Neutrinos

**Hyunmi Song & Jounghun Lee
Seoul National University
(Song & Lee 2011, ApJ, 736, 27)**

MASSIVE NEUTRINOS

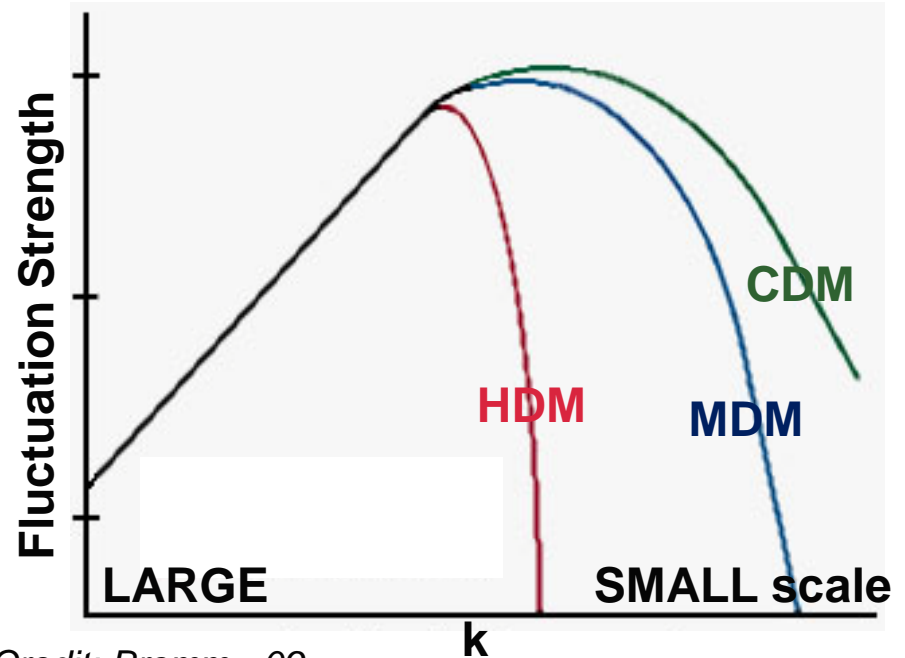
- Solar Neutrino Oscillation reveals neutrinos do have NON-ZERO mass.
- The strongest bounds on neutrino mass comes from Cosmology.
- Mixed Dark Matter = CDM + Massive ν (MDM)



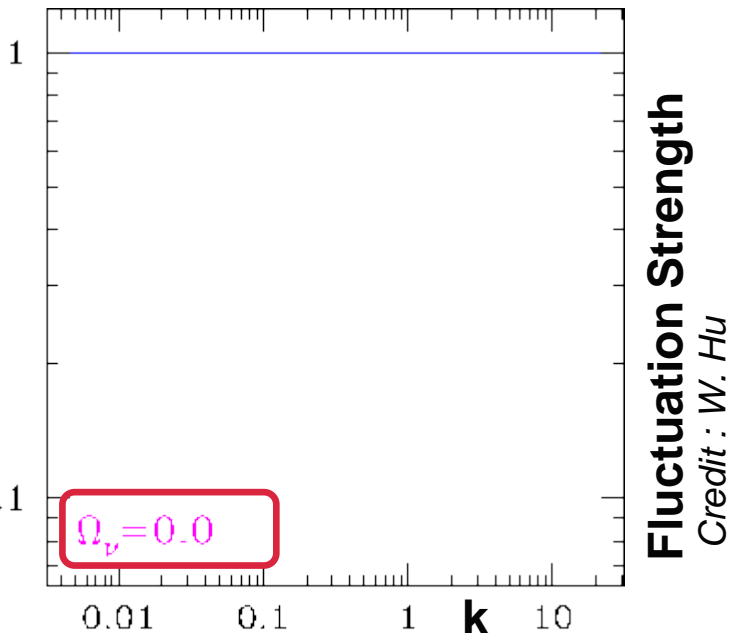
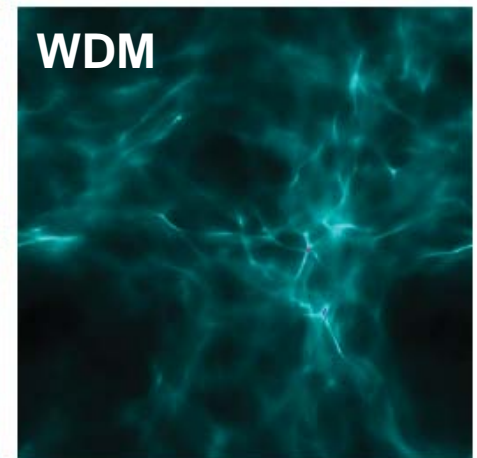
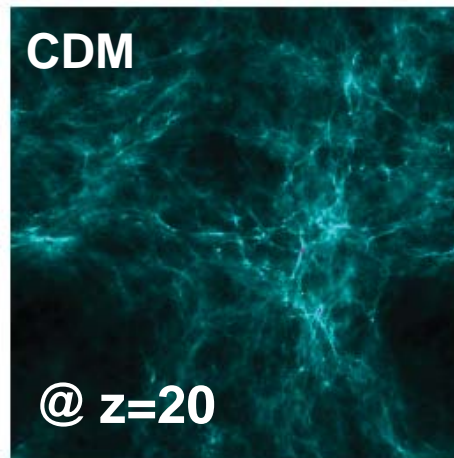
MASSIVE NEUTRINOS

- The growth of fluctuations is suppressed below the free-streaming scale.

EFFECT OF MASSIVE NEUTRINOS

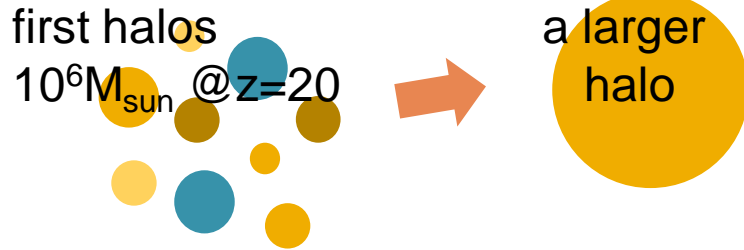


Credit: Bromm+ 09



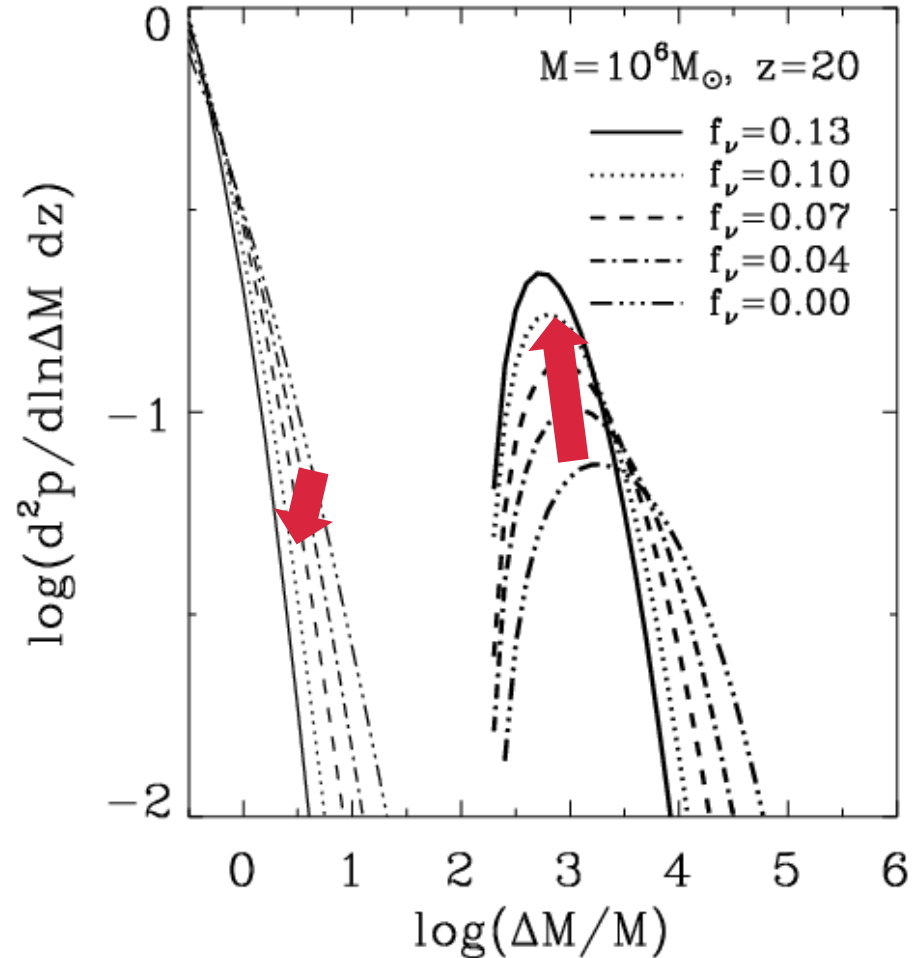
MERGING RATES in MDM

Halo-to-Halo



$$\frac{d^2 P(\text{bigger halo/filament @ } z+dz \mid \text{halo @ } z)}{d \ln \Delta M dz}$$

Halo-to-Filament

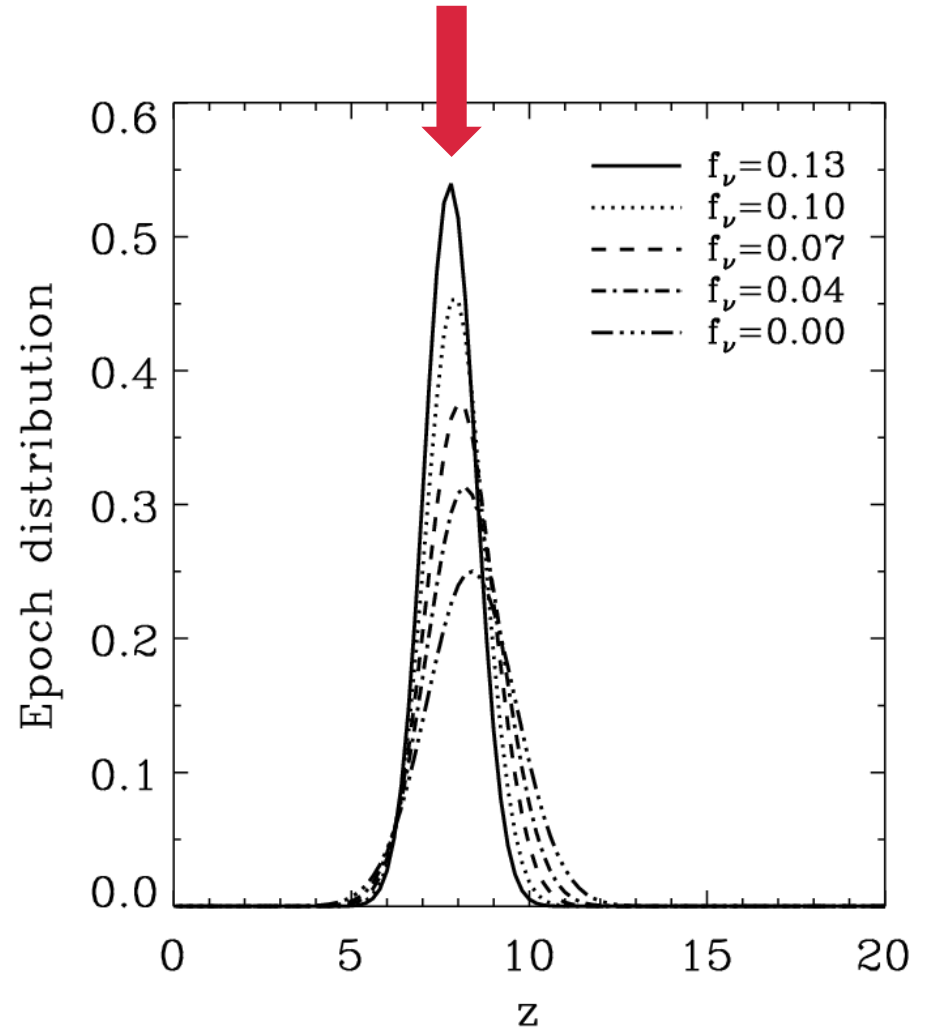
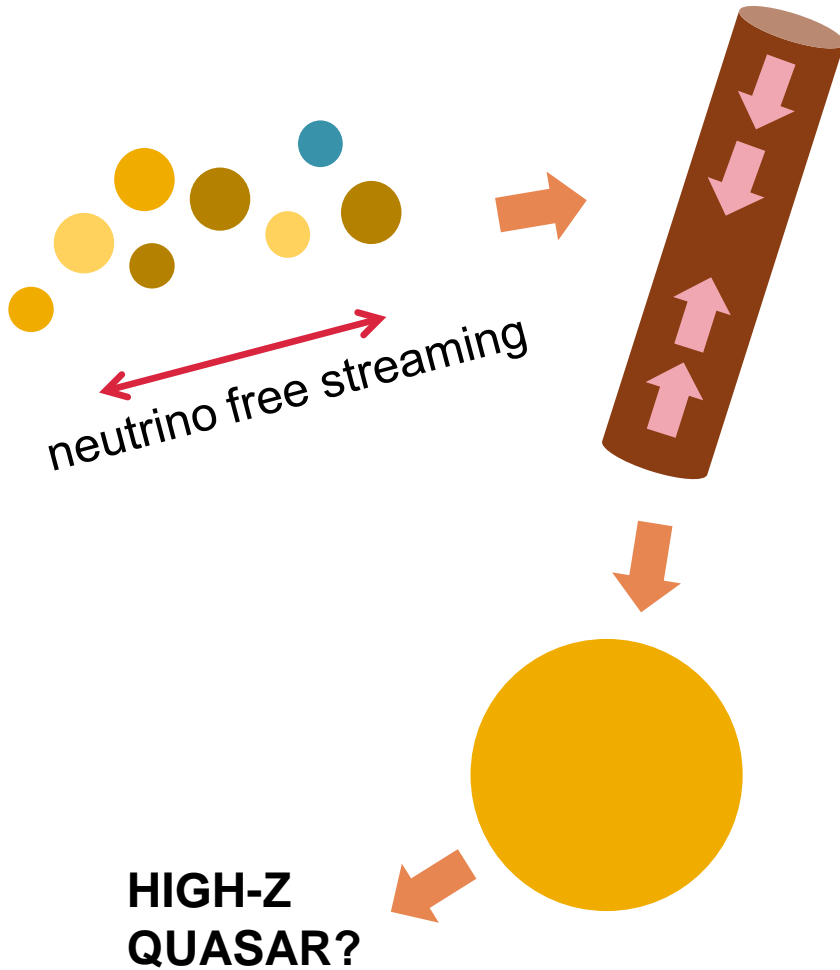


$$f_{\nu} = \Omega_{\nu} / \Omega_m \quad m_{\nu} = (f_{\nu} \Omega_m h^2 / N_{\nu}) 93.2 \text{ eV}$$

thin lines : halo-to-halo

thick lines : halo-to-filament

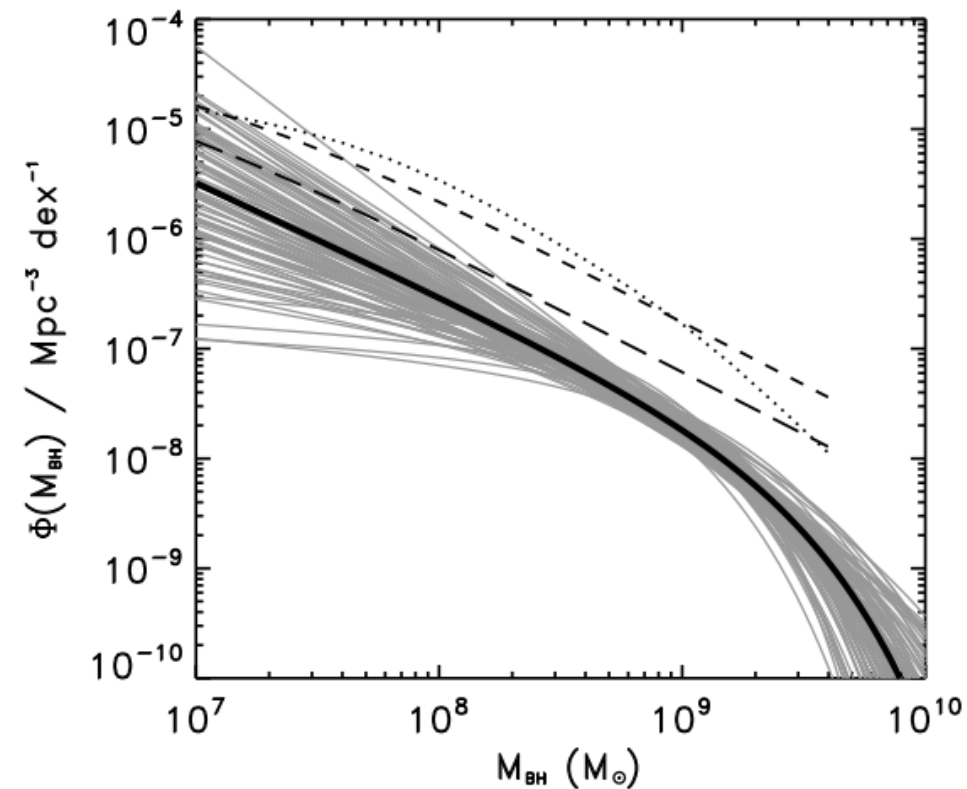
COLLAPSE EPOCH in MDM



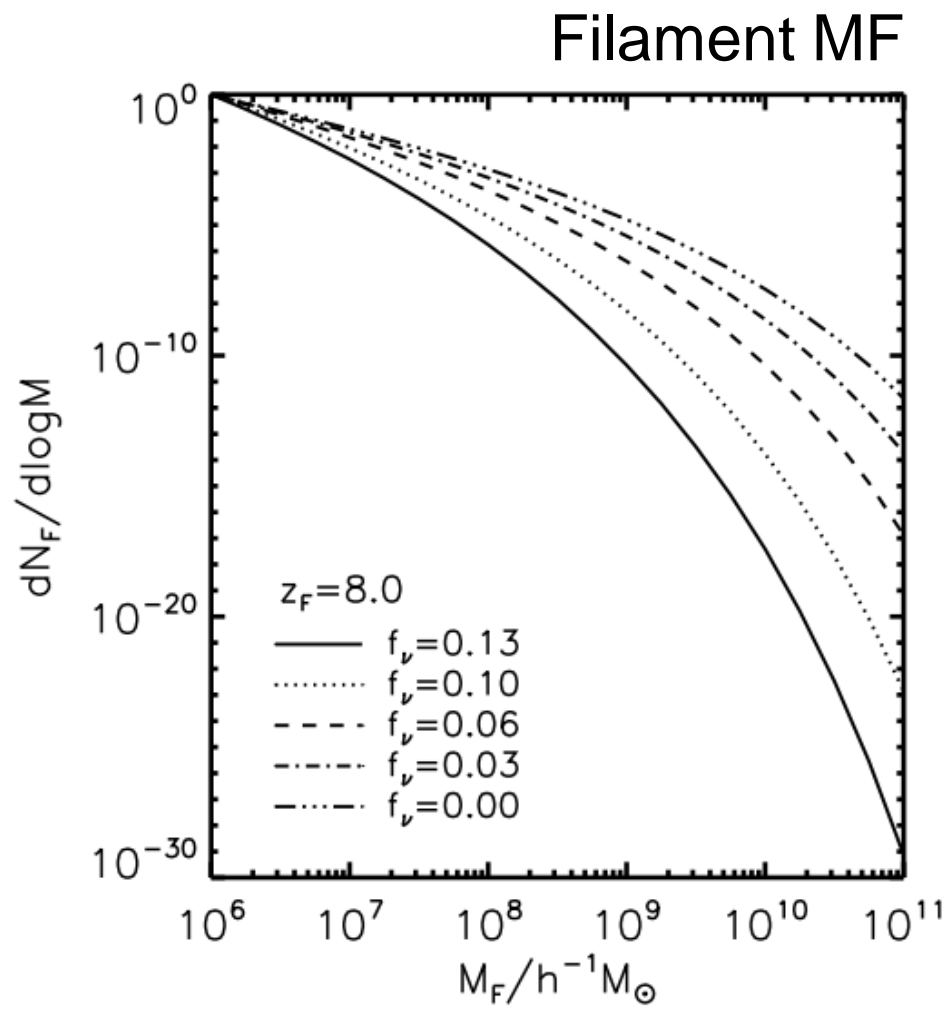
CONNECTION TO HIGH Z QUASARS

- Park & Lee 09
 - Matter and gas accreting more efficiently through filamentary structure. (bridge effect)
- Gao & Theuns 07, Willott + 10
 - Filaments' longest axis collapse would result in forming the first massive galaxies and seeding SMBHs in it.
 - Later these SMBHs would power high-z quasars.
- First Filament MF \Leftrightarrow High-z Quasar MF

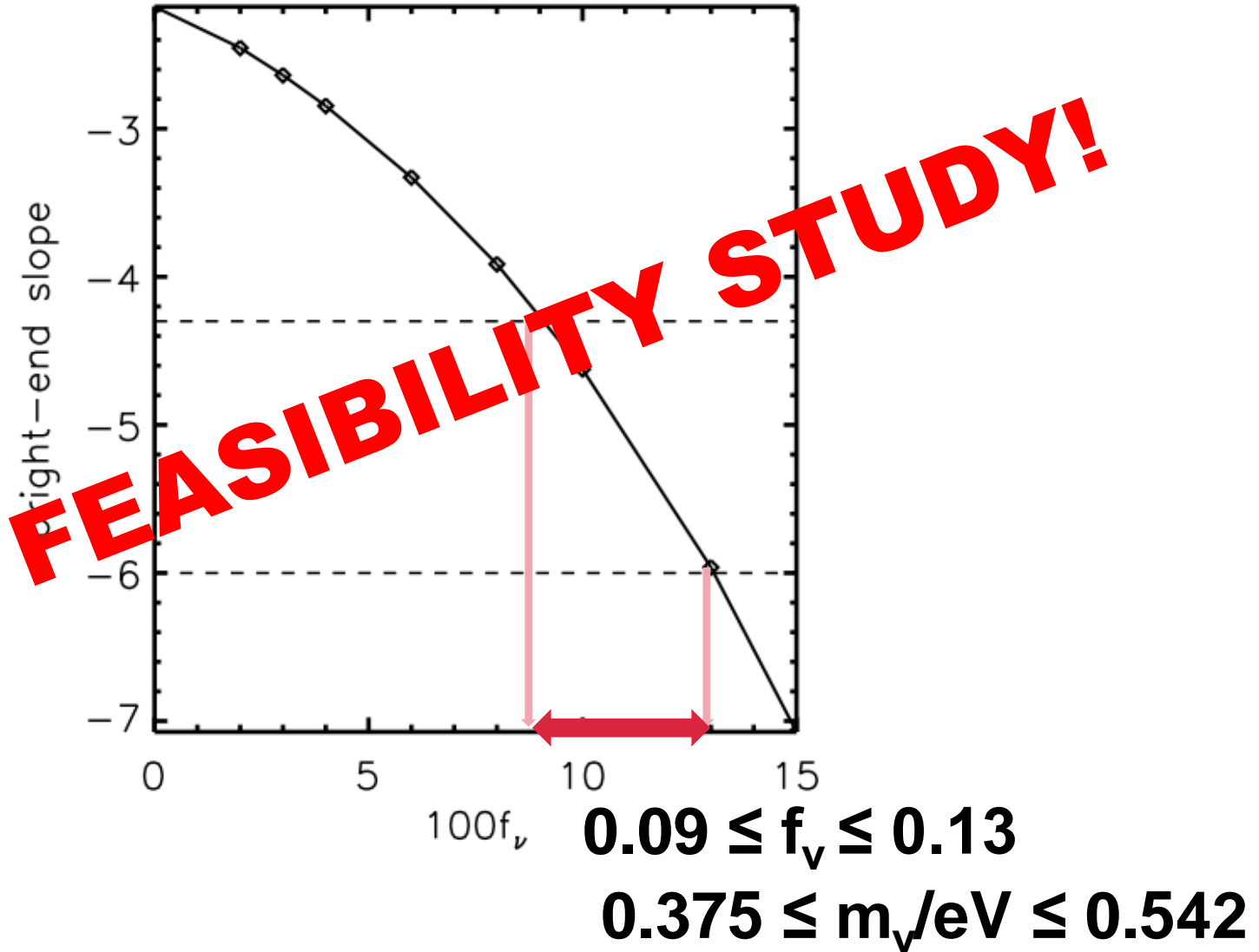
CONSTRAIN NEUTRINO MASS



High-z Quasar MF
Willott + 10



CONSTRAIN NEUTRINO MASS



SUMMARY

- We explore the effect of massive neutrinos on the formation and evolution of the first filaments.
- We suggest that the first filaments may open a new channel to form high- z quasars and we also test the possibility of constraining the neutrino mass from the high- z quasar mass function.
- Recently, for another approach to constrain neutrino mass, we examine the alignment of cluster galaxies. (arXiv:1106.5104)

Thank You 😊