

Testing CLES 18

Task 1.5

Josefina Montalban & Yveline Lebreton

Comparison models, Aarhus 24-28 October 2005

Task1.5

PARAMETERS:

mass	2.0 M _⊙
Z	0.02
X	0.72
α	1.60
ov	0.15 Hp
Xc	0.01

Isotopic ratios

	CLES	CESAM
² H/H	3.4001 10 ⁻⁵	3.010 10 ⁻⁵
³ He/ ⁴ He	1.420 10 ⁻⁴	1.100 10 ⁻⁴
		4.185 10 ⁻⁴

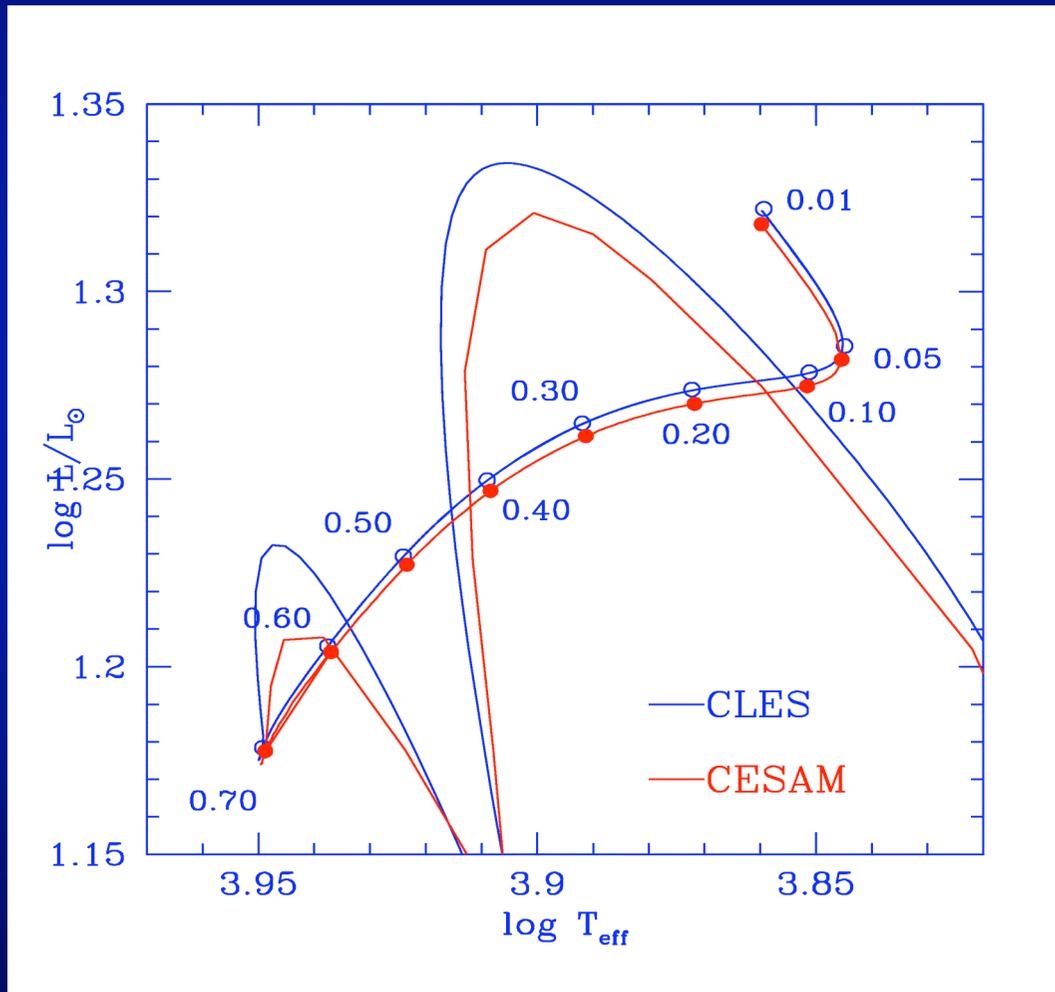
PHYSICS:

Conv.	MLT (Bohm-Vitense 68 + Heyney et al. 65)
Atm.	Gray Eddington
Opacities	OPAL (96) + Alexander & Ferguson(94)
EOS	OPAL (2001, Tables)
Reaction rates	NACRE (Angulo et al. 9)
Mixture	Solar (Grevesse & Noels)
Overshoot	$\nabla = \nabla_{ad}$ (Zahn 91)

GN93 mixture

	CLES	CESAM
He	10.93	10.99
Li	3.31	1.16
Be	1.42	1.15
B	2.60	2.60

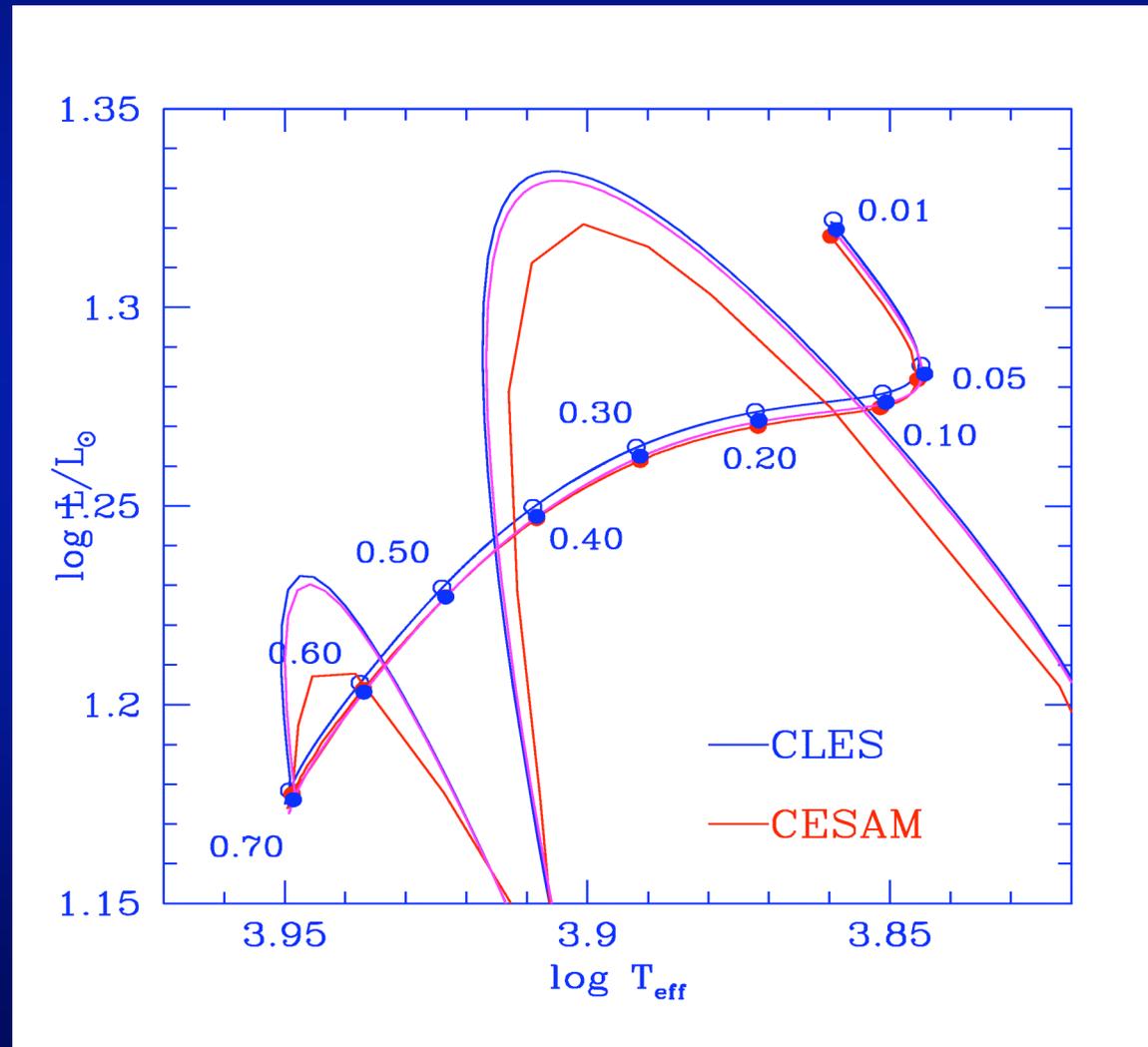
HR diagram



$\Delta T_{\text{eff}} < 0.2\%$

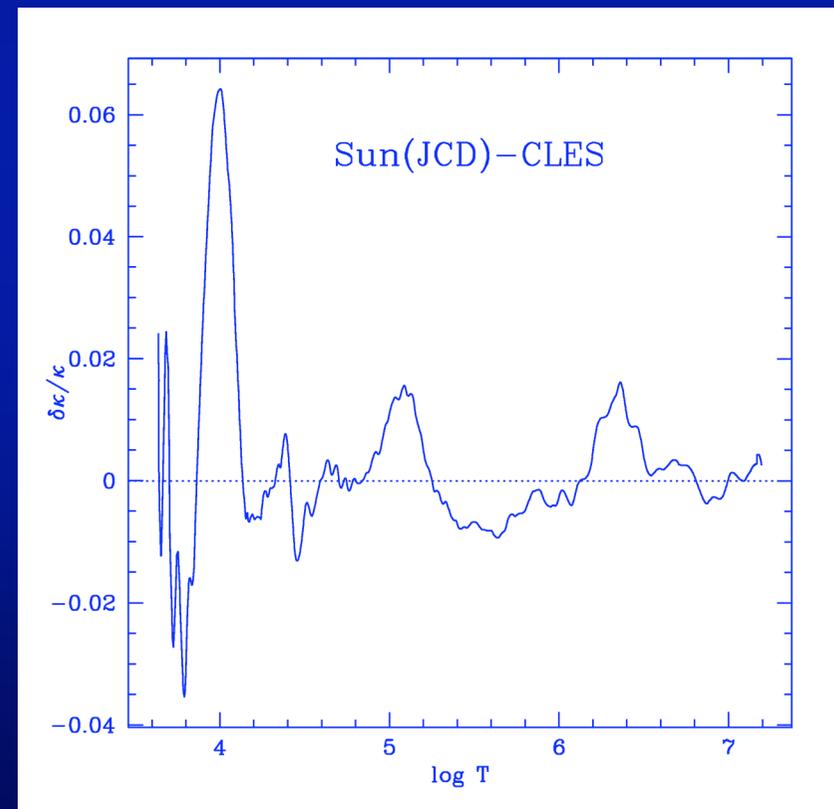
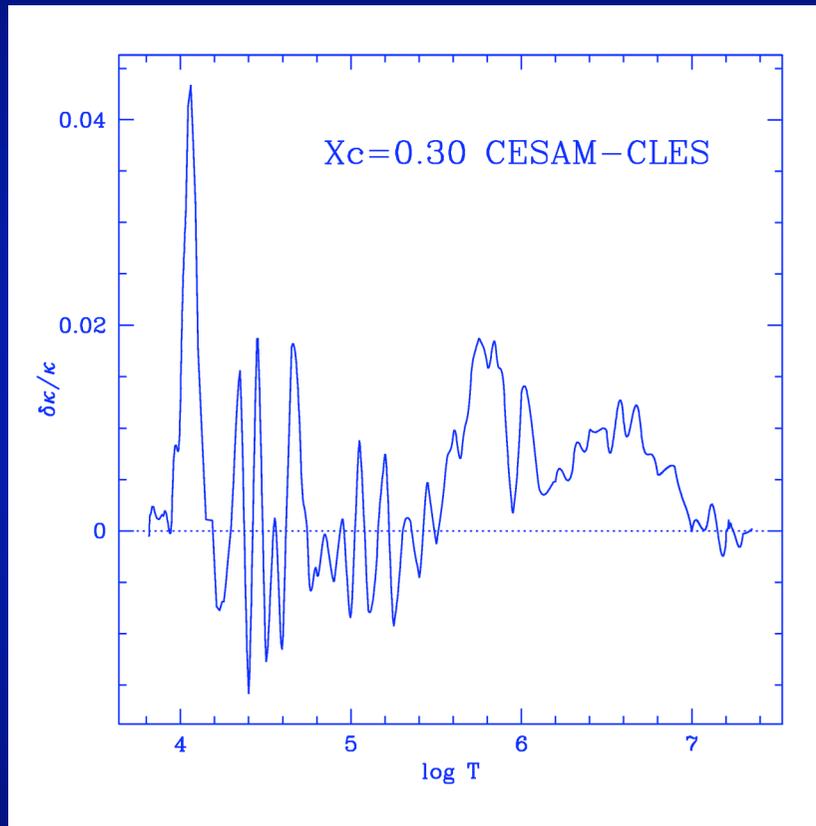
$\Delta L/L_{\odot} < 1\%$

HR diagram:

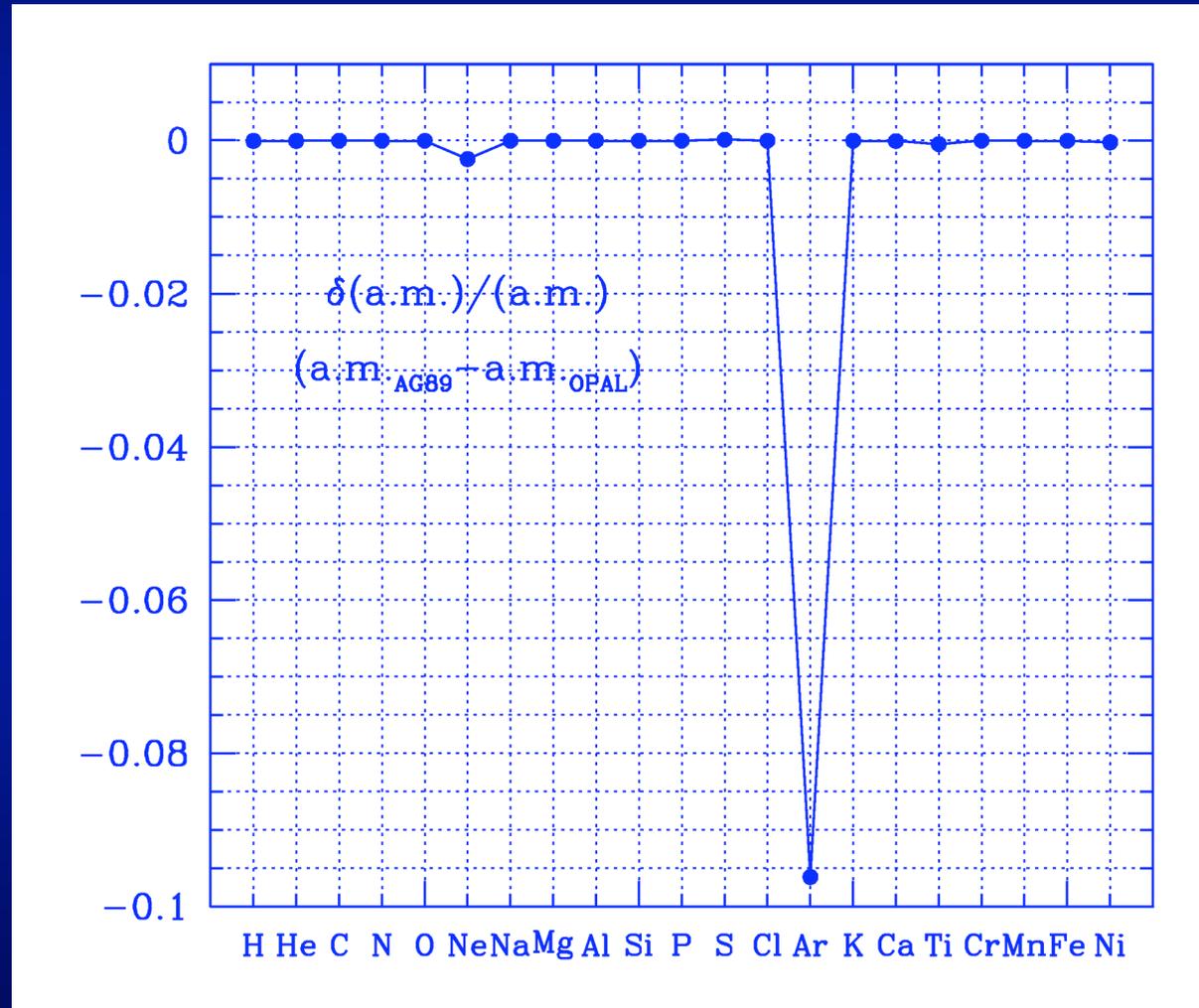


K_{CLES}
increased by
0.5%

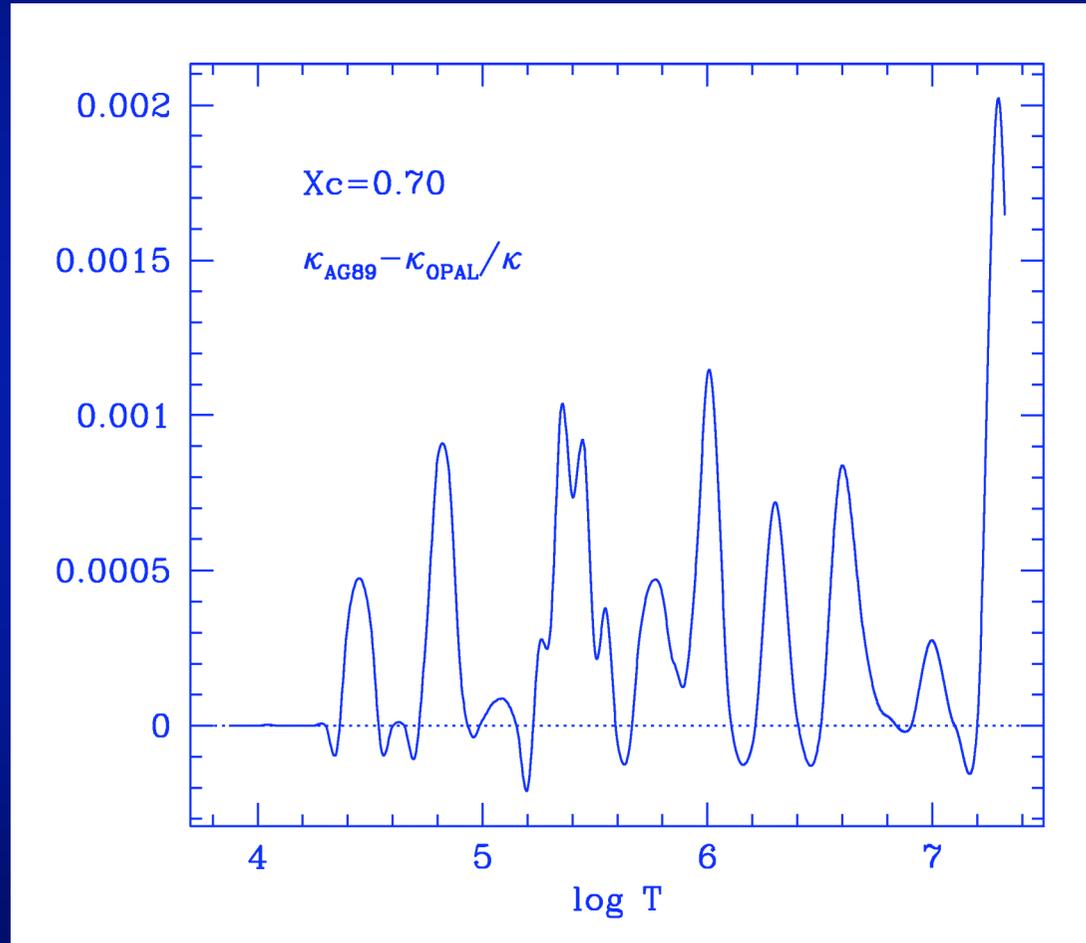
Opacity



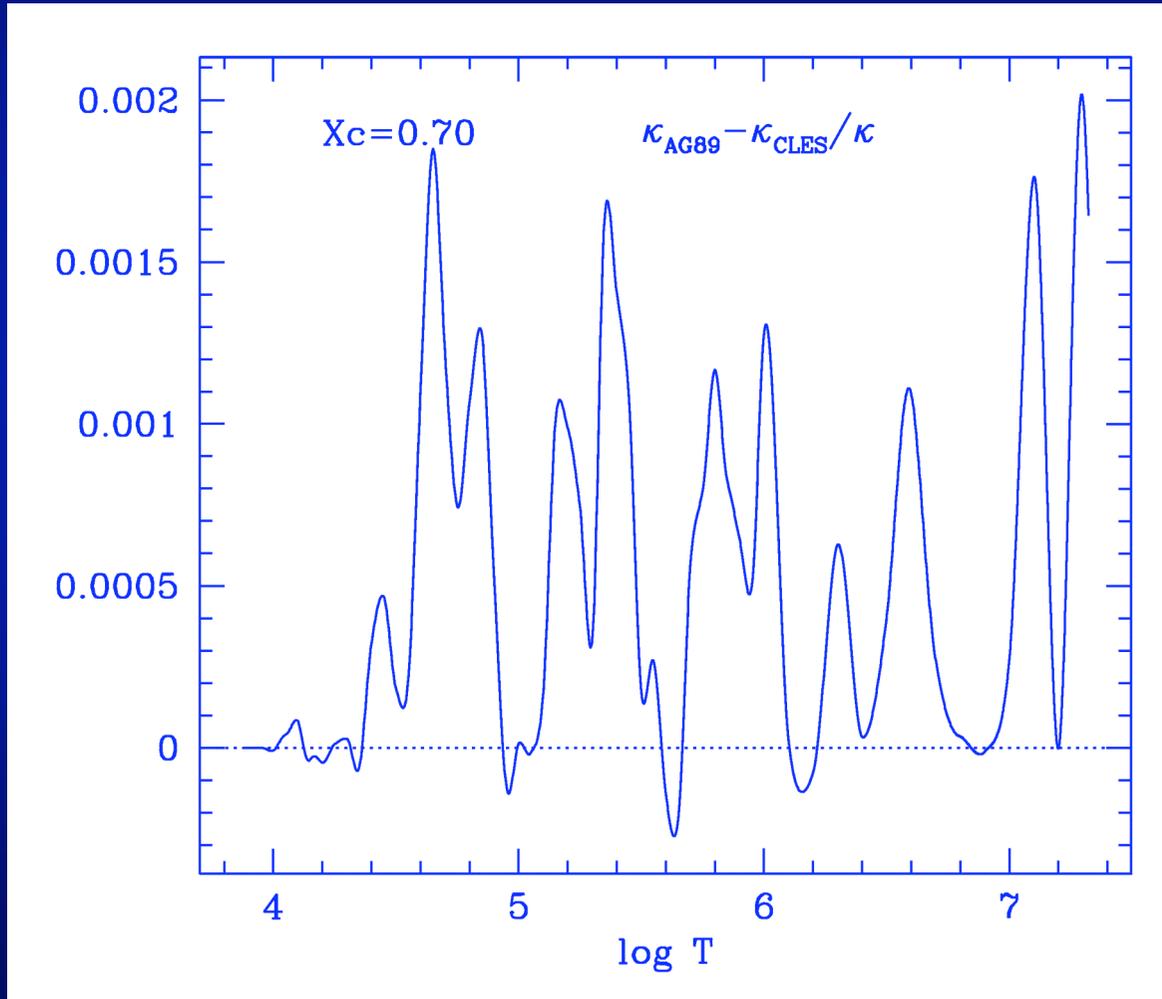
Opacity : effect of atomic mass



Opacity : effect of atomic mass



Opacity : GN93 ??



$$X(F) = 2.5 \cdot 10^{-5} \left\{ \begin{array}{l} O \\ Ne \end{array} \right.$$

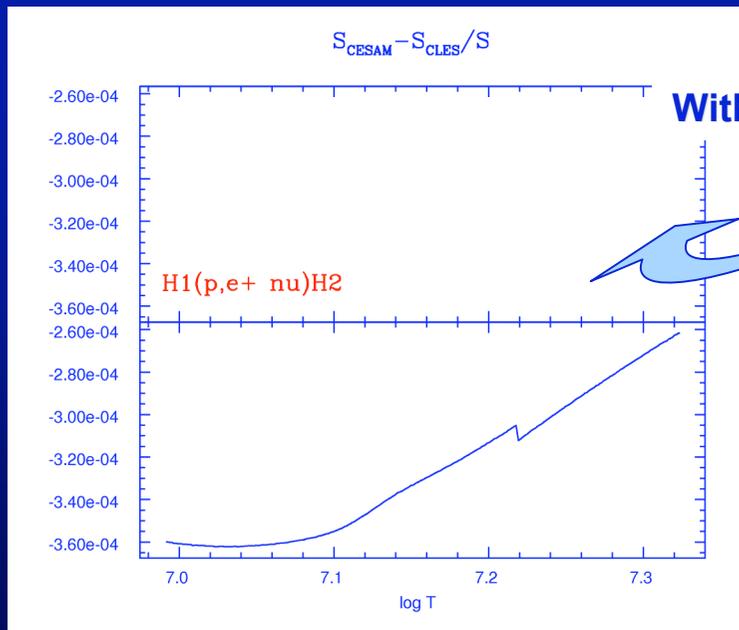
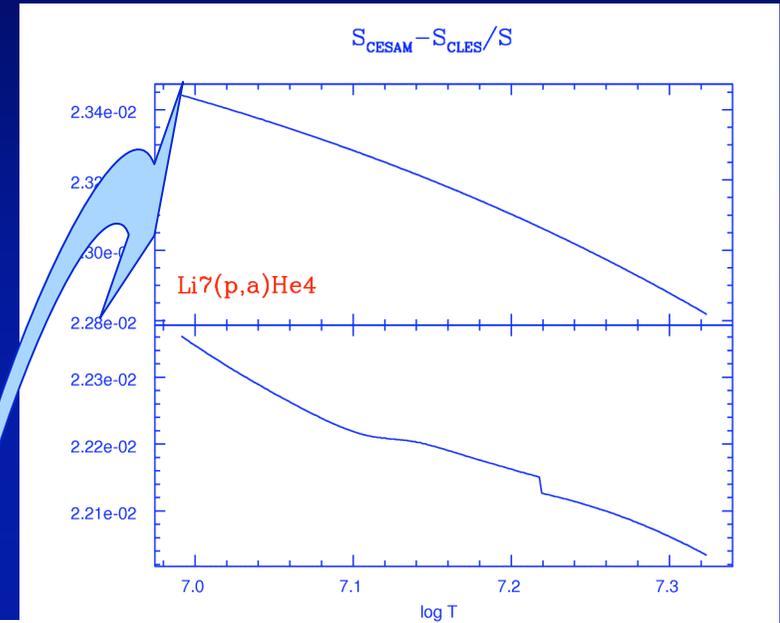
$$X(Sc) = 1. \cdot 10^{-6} \left\{ \begin{array}{l} Ca \\ Ti \end{array} \right.$$

$$X(V) = 7. \cdot 10^{-6} \left\{ \begin{array}{l} Ti \\ Cr \end{array} \right.$$

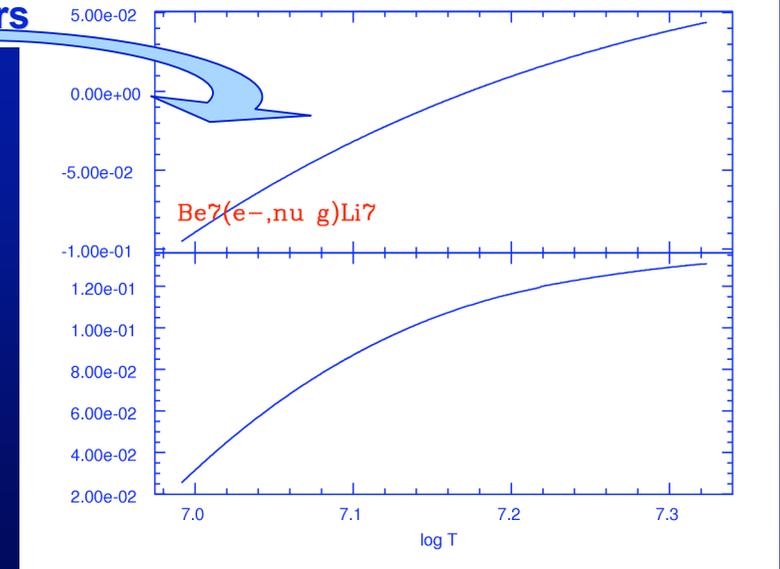
$$X(Co) = 5.8 \cdot 10^{-5} \left\{ \begin{array}{l} Fe \\ Ni \end{array} \right.$$

Nuclear reaction rates: PP

1. $\text{H}(p, e^+, n)^2\text{H}$	-3. e-4	< e-8
2. $^2\text{H}(p, g)^3\text{He}$	-3. e-4	1.e-7
3. $^3\text{He}(^3\text{He}, 2p)^4\text{He}$	-1.5e-3	< e-8
4. $^3\text{He}(a, g)^7\text{Be}$	-1.5e-3	< e-8
5. $^7\text{Li}(p, a)^4\text{He}$	2. e-2	1.3e-2
6. $^7\text{Be}(e^-, n g)^7\text{Li}$	1.5e-2	-0.1, 5.e-2
7. $^7\text{Be}(p, g)^8\text{B}(, e^+ n)^8\text{Be}(a)^4\text{He}$	-1.5e-3	< e-8

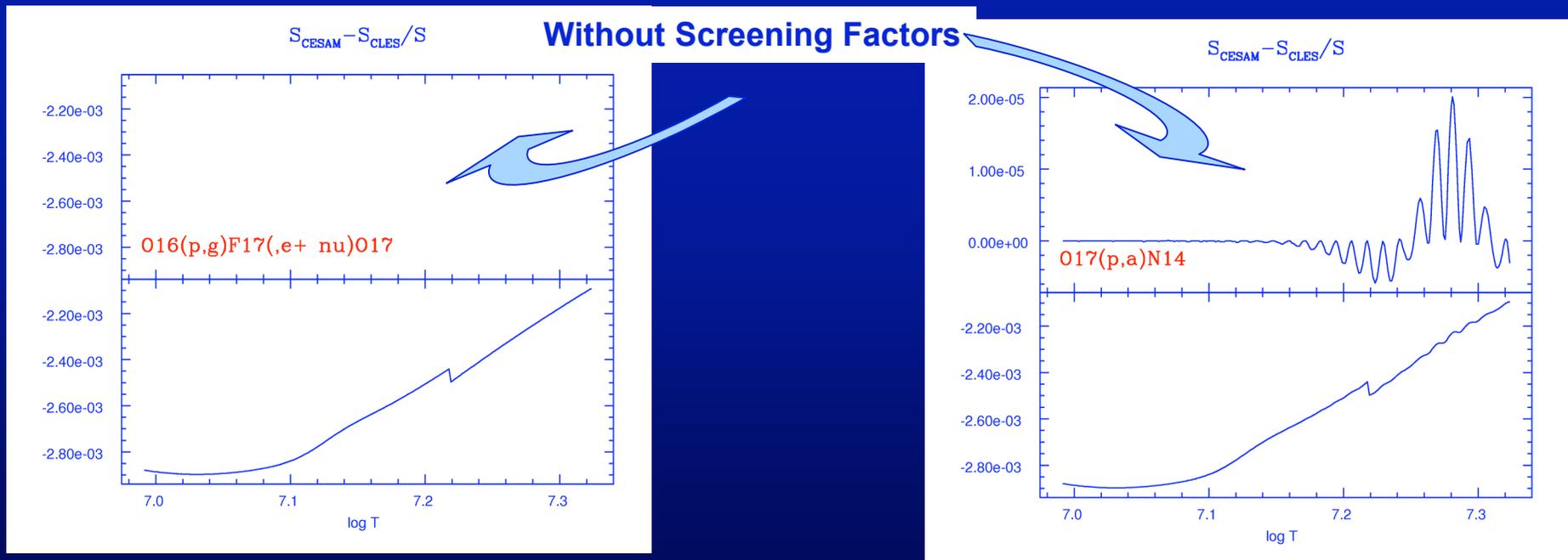


Without Screening Factors



Nuclear reaction rates: CNO

1. $^{12}\text{C}(p,g)^{13}\text{N}(e^+)^{13}\text{C}$	-2. e-3	< e-8
2. $^{13}\text{C}(p,g)^{14}\text{N}$	-2. e-3	< e-8
3. $^{14}\text{N}(p,g)^{15}\text{O}(e^+,n)^{15}\text{N}$	-2.2e-3	< e-8
4. $^{15}\text{N}(p,g)^{16}\text{O}$	-2.2e-3	< e-8
5. $^{15}\text{N}(p,a)^{12}\text{C}$	-2.2e-3	1.e-6
6. $^{16}\text{O}(p,g)^{17}\text{F}(e^+,n)^{17}\text{O}$	-2.5e-3	< e-8
7. $^{17}\text{O}(p,a)^{14}\text{N}$	-2.5e-3	2.e-5(max)



Screen-Factor: (Salpeter 1954)

CESAM:

$$f = \exp \left(1.88 d^8 z_1 z_2 \left(\frac{\rho \zeta}{T^3} \right)^{\frac{1}{2}} \right)$$

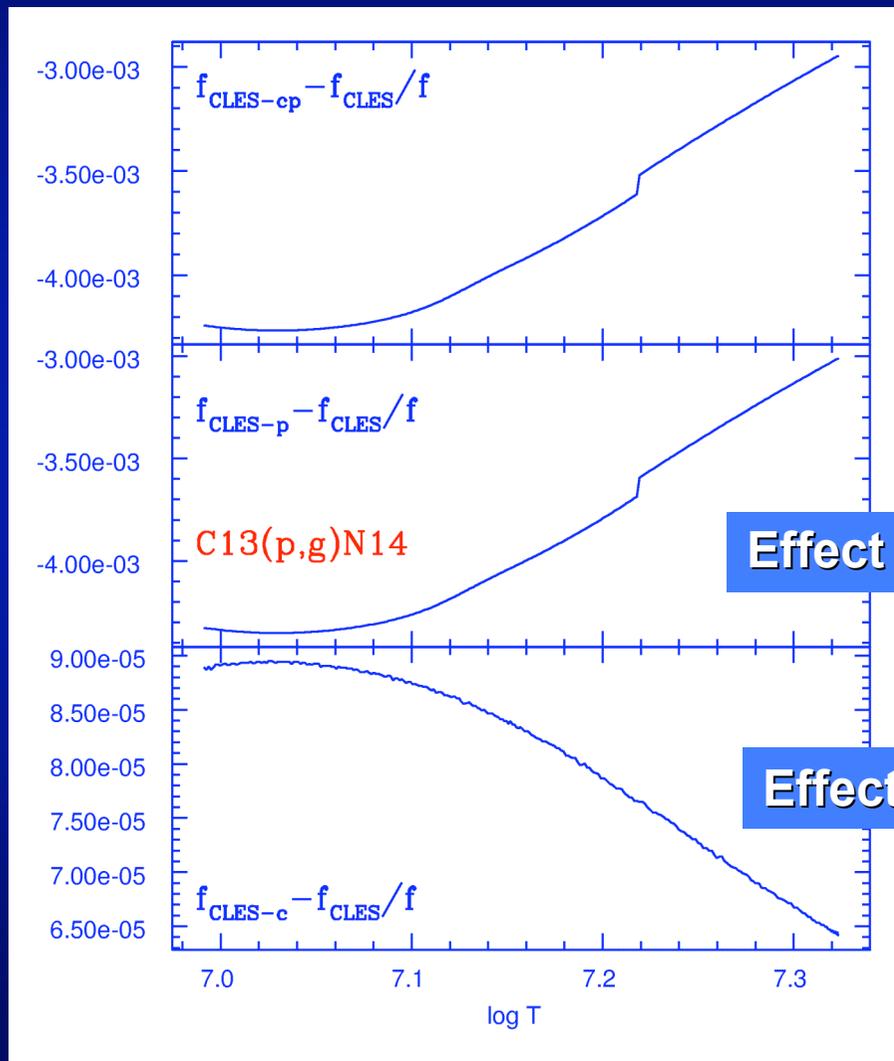
$$\zeta = \sum_i z_i (1 + z_i) x_i$$

CLES:

$$f = \exp \left(5.942168 d^{-6} z_1 z_2 \left(\frac{\rho \psi}{T_9^3} \right)^{\frac{1}{2}} \right)$$

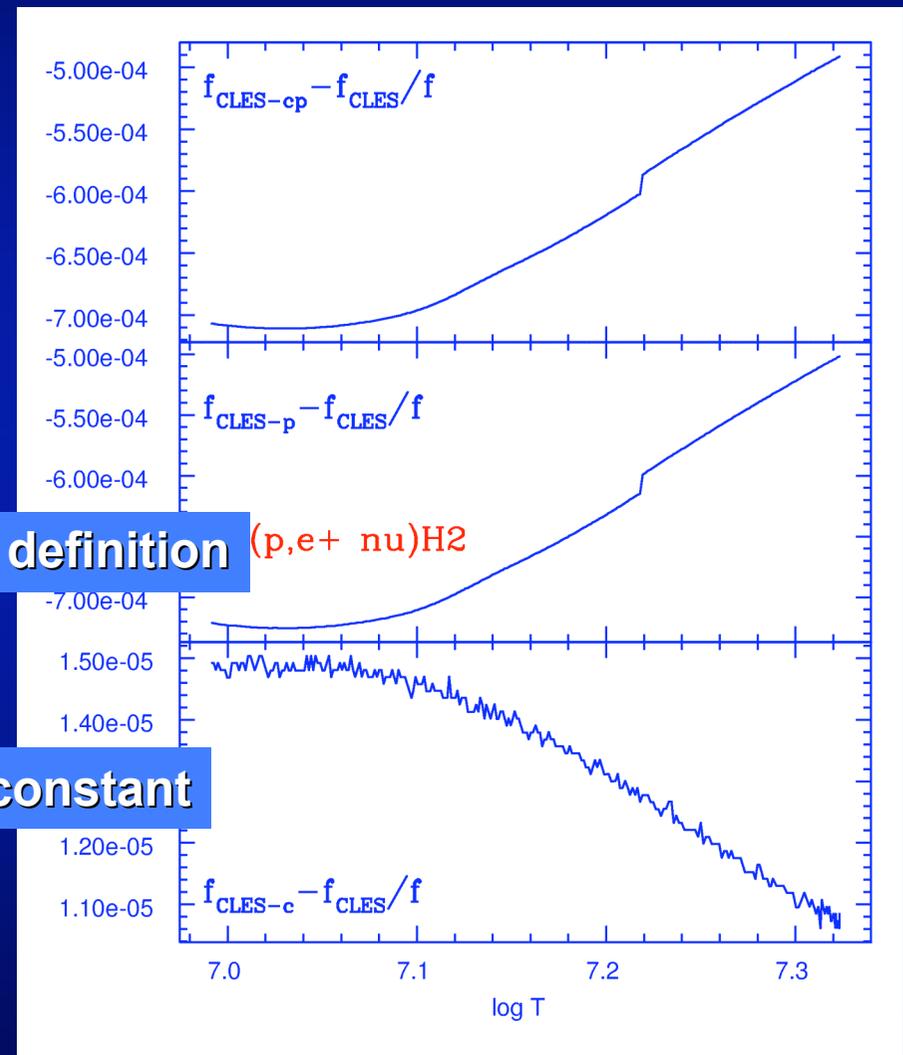
$$\zeta = \sum_{1-4} z_i (1 + z_i) x_i + Z(1 + Z)x(Z)$$

Screen-Factor:

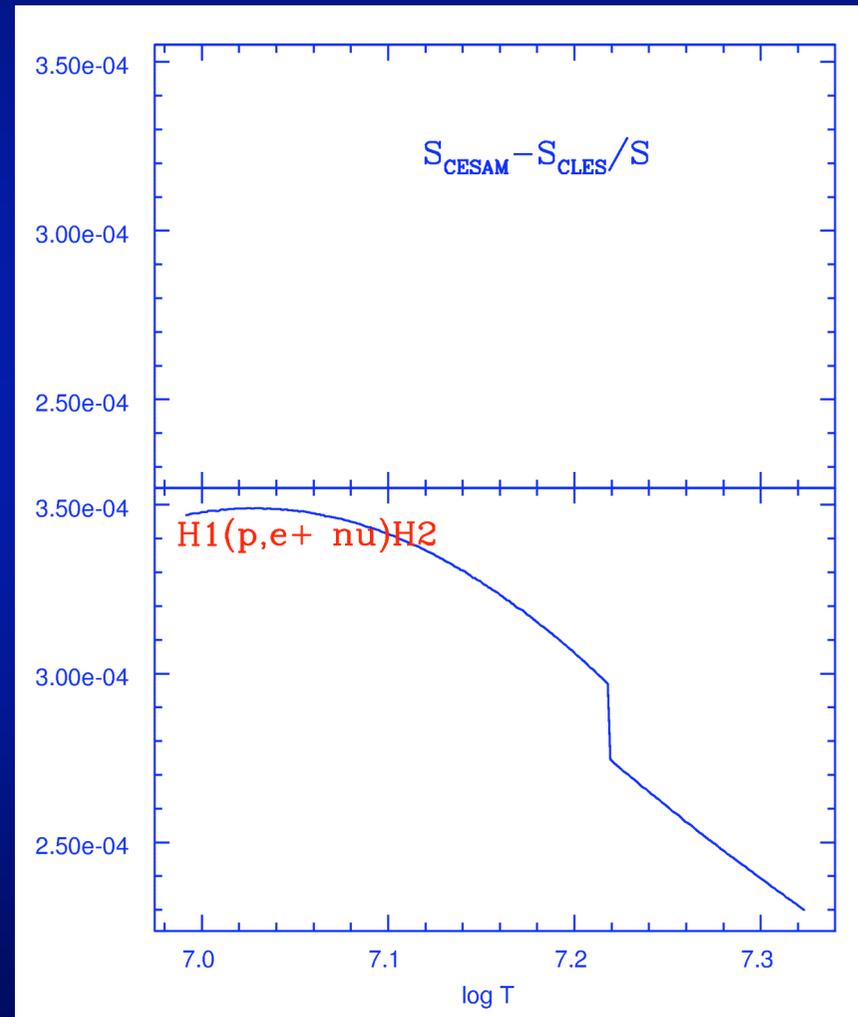
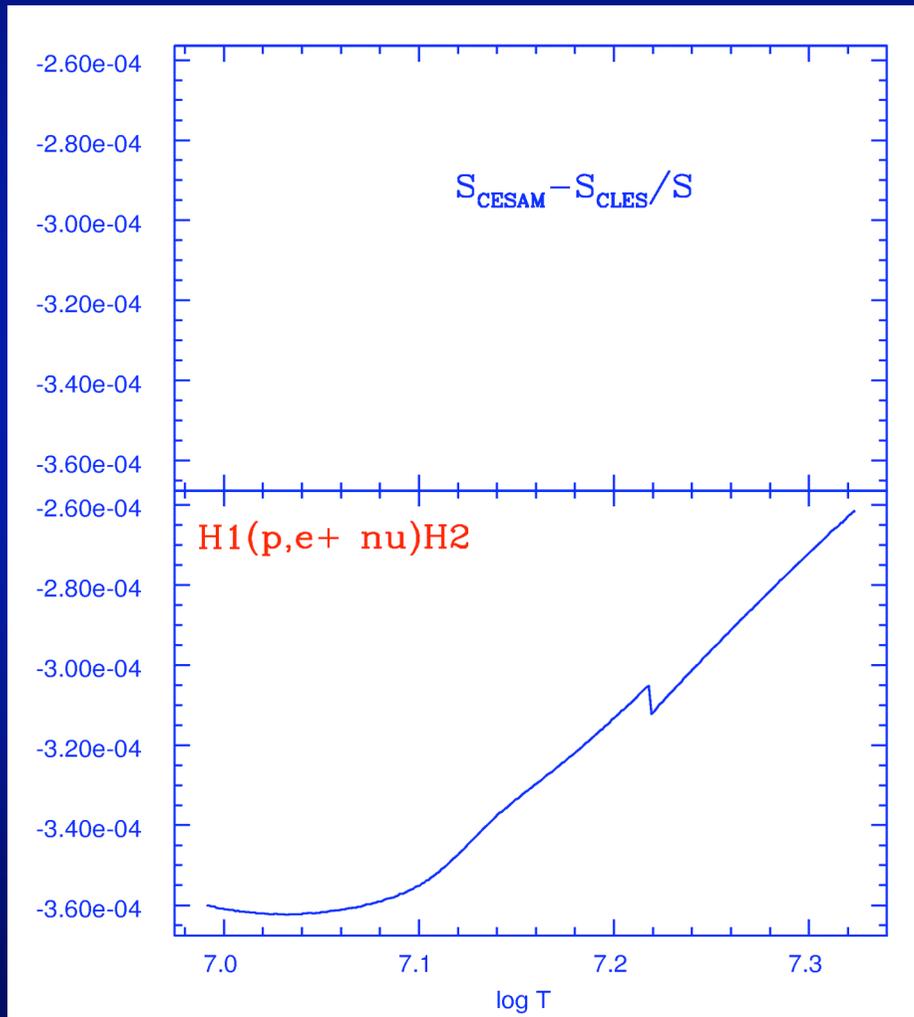


Effect of ζ definition (p,e+ nu)H2

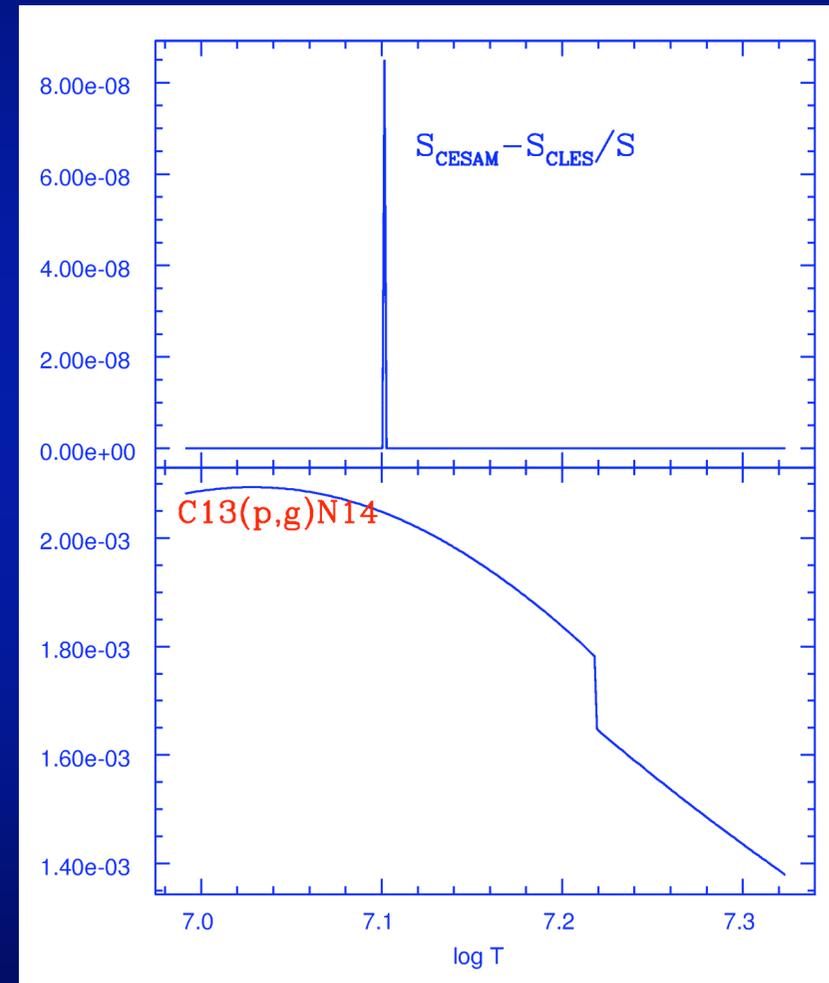
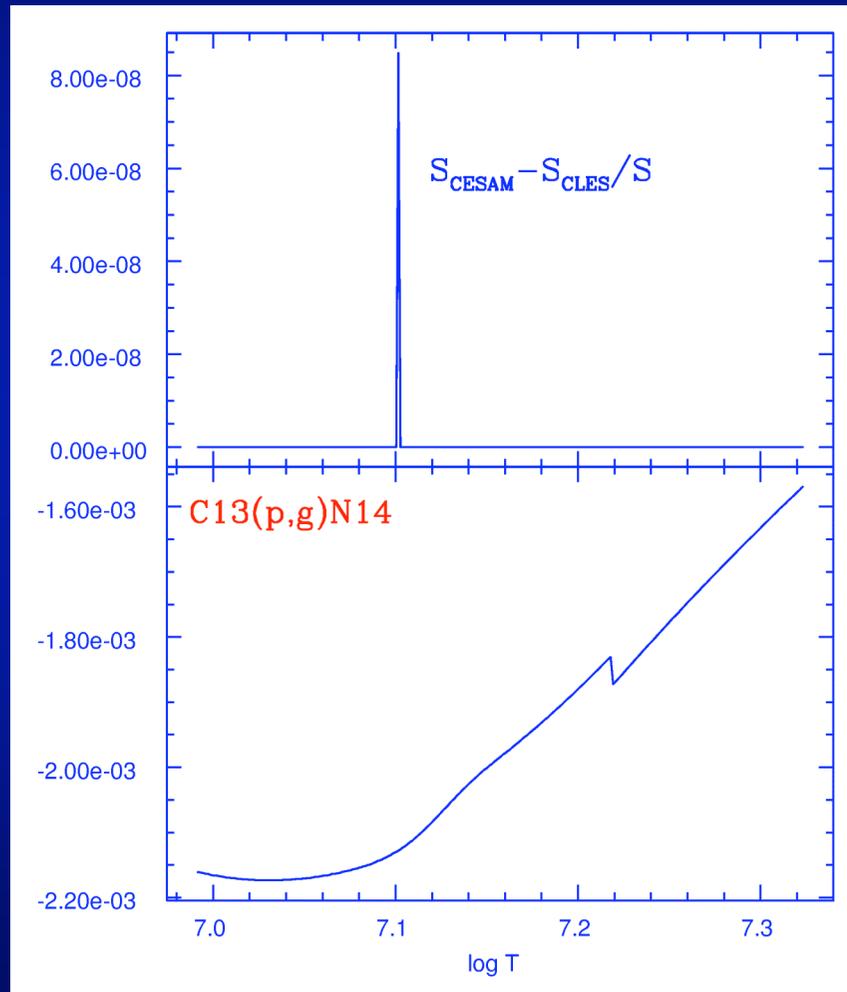
Effect of constant



Nuclear reaction rates: PP



Nuclear reaction rates: CNO



Summary

1. Opacity: CESAM > CLES , average: 1%, more than 4% in surface

2. EoS :

1. Interpolation:

Γ_1 :	0.2 %	$T < 10^5$
C_p :	0.6 %	$T < 10^5$
	0.2 %	$T > 10^7$
δ :	$5 \cdot 10^{-4}$	$T < 10^5$
P :	$6 \cdot 10^{-4}$	$T < 10^5$
	$2 \cdot 10^{-4}$	$T > 10^7$
∇_{ad} :	0.5%	$T < 10^5$

2. Inconsistency

Γ_1 :	1.5 %	$T < 10^5$
C_p :	5 %	$T < 10^5$
δ :	$2 \cdot 10^{-6}$	numeric
P :	$\pm 10^{-7}$	numeric
∇_{ad} :	2 – 4 %	$T < 10^5$

Summary

1. Nuclear reaction rates: Main difference is in Screen Factors
2. Convective core:
 1. Difference in convective core mass $\Delta m_{cc}/M = 6 \cdot 10^{-4}$
 2. Difference in N_{BV} profile ... a suivre...