

CESAM2K

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Workshop

Numerics

- Solution of the stellar structure eqs. by collocation method based on piecewise polynomial approximations projected on their B-spline basis
 - Stable and robust calculations
 - Restitution of the solution not only at grid points
 - Automatic mesh refinement
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Numerics

- Precise restoration of the atmosphere
 - Modular in design
 - Evolution of the chemical composition:
 - Without diffusion: implicit Runge-Kutta scheme
 - With diffusion: solution of the diffusion eq. using the Galerkin method
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Equation of state

- OPAL 2001 EoS tables used with the OPAL interpolation scheme
- Other EoS available:
 - EFF
 - CEFF
 - MHD



Opacity

- OPAL 1996 + Alexander & Ferguson 1994:
 - Houdek9 interpolation
 - Interpolation by Lagrange Polynomials for $\log T_6$ and $\log R$, quadratic interpolation for X and Z
 - Other opacities:
 - Interpolation of the tables of Livermore 1991
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Energy generation rate

- NACRE compilation (other compilations available, e. g. Caughlan and Fowler 1988).
 - Several networks available (PP+CNO+3 α). Examples:
 - PP+CNO chains with ^2H and ^7Be in equilibrium
 - PP+CNO without any element in equilibrium
 - PP+CNO+3 α
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Convection

- Böhm-Vitense MLT taking into account the optical thickness of the bubble.
 $l = \alpha H_p$
- *CM (1991) with $l = \alpha H_p$*
- *CGM (1996)*

Atmosphere

- $T(\tau)$ relation based on the grey Hopf atmospheric law. Connection with the envelope at $\tau=10$.
 - Other $T(\tau)$ relations available:
 - Eddington law
 - $T(\tau)$ for the Sun calculated with Atlas 9 of Kurucz (for $T_{\text{eff}}=5750, 5777$ K)
 - $T(\tau)$ calculated with Atlas 12 of Kurucz (for $[Fe/H]=0, -0.2, -0.5, -1.0$)
 - Single shell
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Detailed description of the code

• Code available upon request at

http://www.obs-nice.fr/morel/CESAM_PROTECT/CESAM.html

• Package contains a detailed description of the code (in french), as well as a user's guide

• Description in english: Morel 1997, A&AS, 124, 597

CESAM2K V2

- Many mesh points added around the CZ/RZ limits to refine the profile of the Brunt-Väissälä frequency
- Possibility to use either the radiative or the adiabatic gradient on the overshoot zones
- Diffusion of the angular momentum

