Tests and future plans for the FRANEC code (Pise and Naples groups)

1) Deeper comparison of the case 1.4

Parameters:

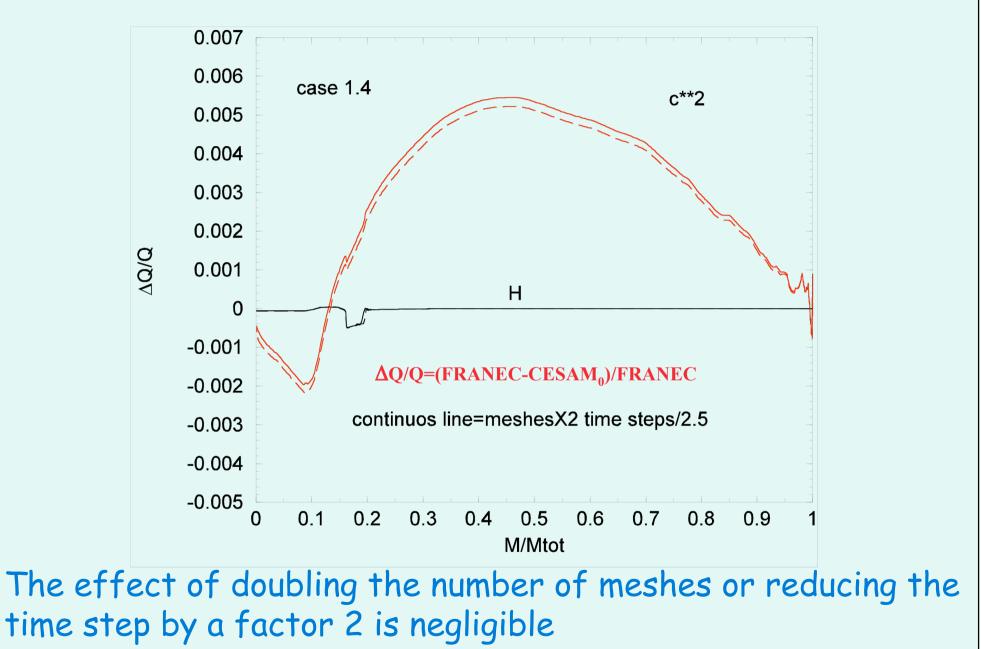
 $M/M_0=2.0$ $X_0=0.70$ $Y_0=0.28$ $Z_0=0.02$ $I_{MLT}/H_p=1.6$

Target model: $T_c = 1.9 \cdot 10^7 \text{ K} (PMS)$

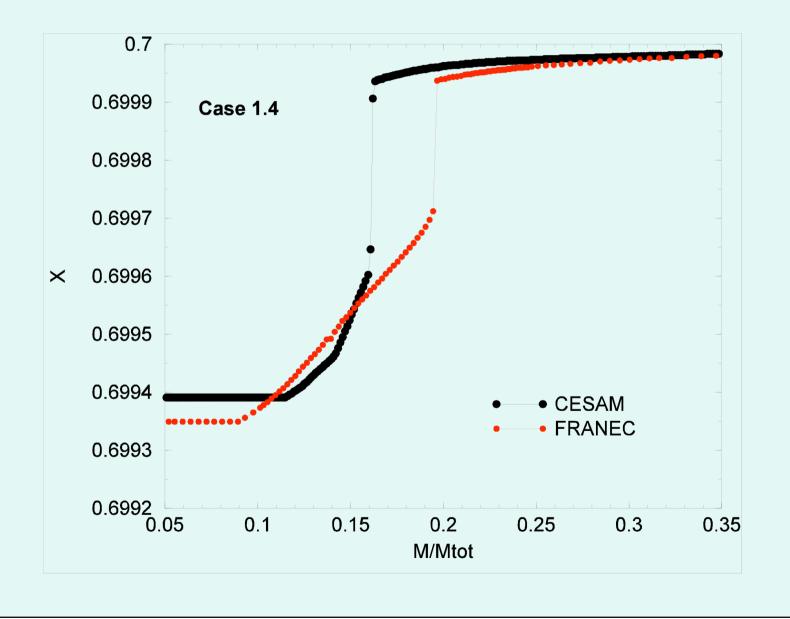
Global parameters case 1.4:

	CESAMo	FRANEC	Δ
•Age [Myr]	7.043	7.764	0.10
·R/Ro	1.866	1.878	0.006
·L/Lo	15.80	16.31	0.030
•T _{eff} [K]	8431	8472	0.005
•T _c [10 ⁷ K]	1.900	1.900	0.
• ρ_c [g/cm ³]	49.22	50.47	0.025
٠X _c	0.6994	0.6993	-0.0001
•M _{cor} /M	0.1075	0.0945	-0.121
•R _{env} /R	0.9988	0.9998	0.001

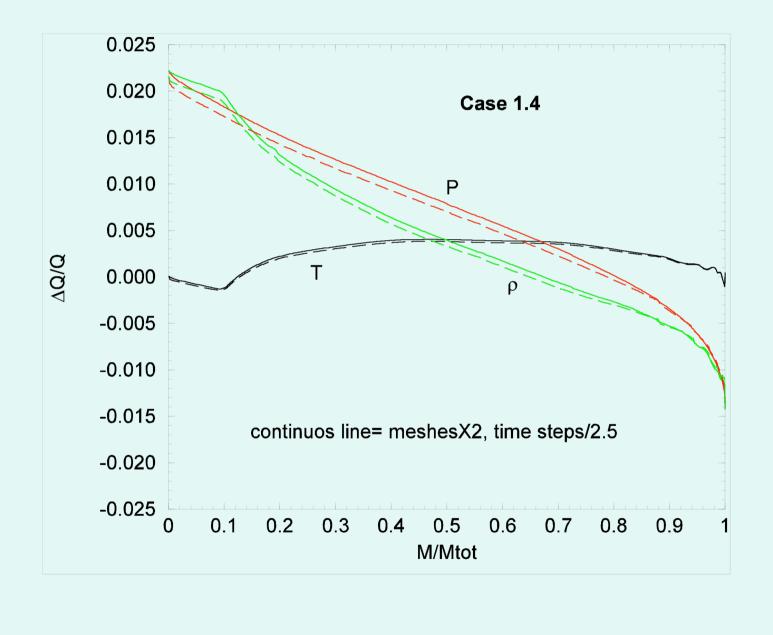
Relative differences for c^2 and X abundance



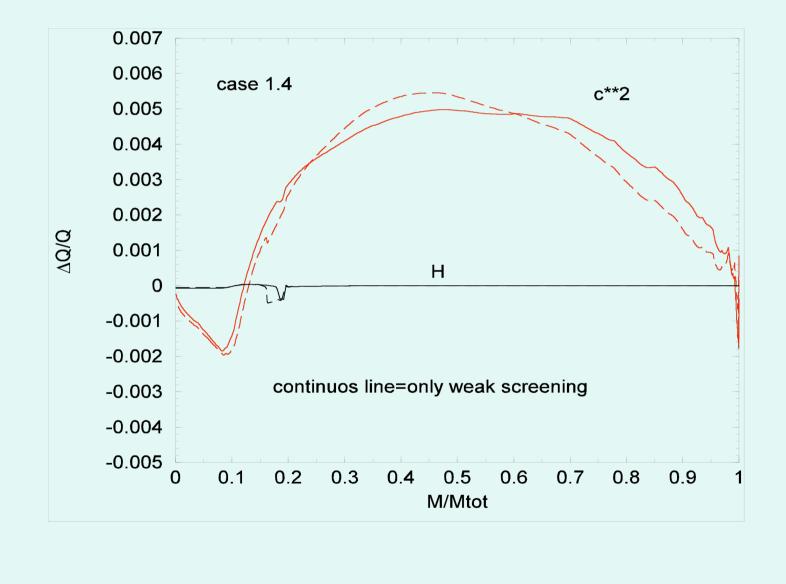
Comparison of the H profile



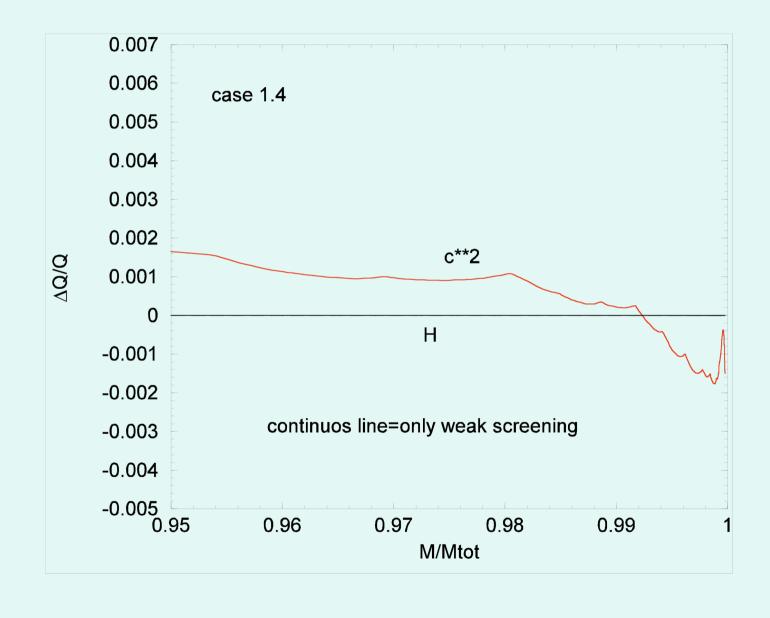
Relative differences for T, P, ρ

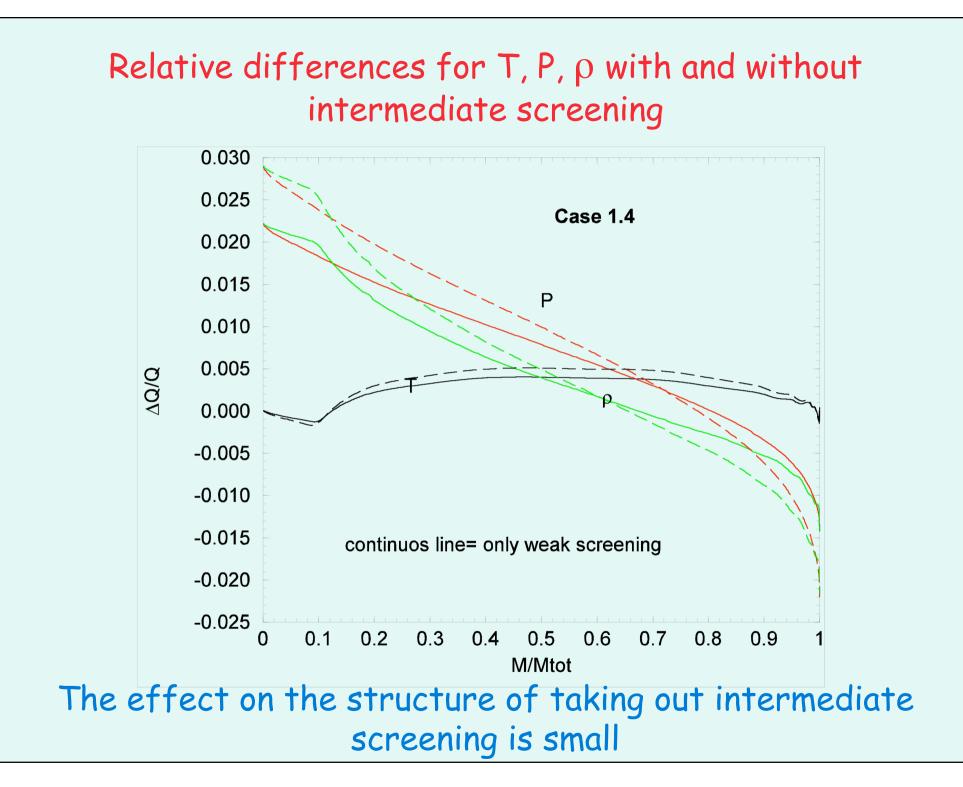


Only weak (Salpeter) screening for a better comparison with $\ensuremath{\mathsf{CESAM}}_0$

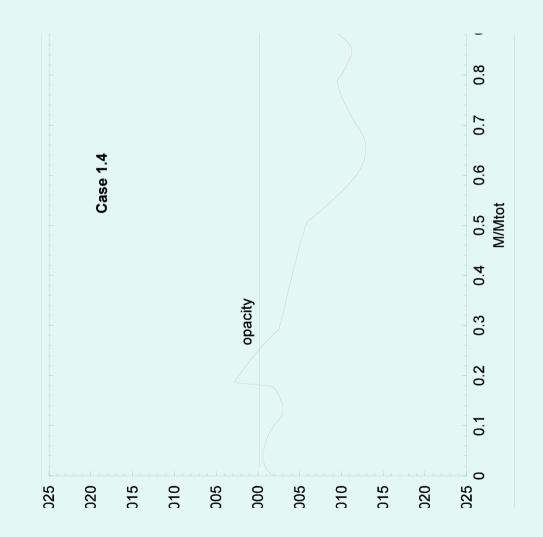


Zoom of the more external region

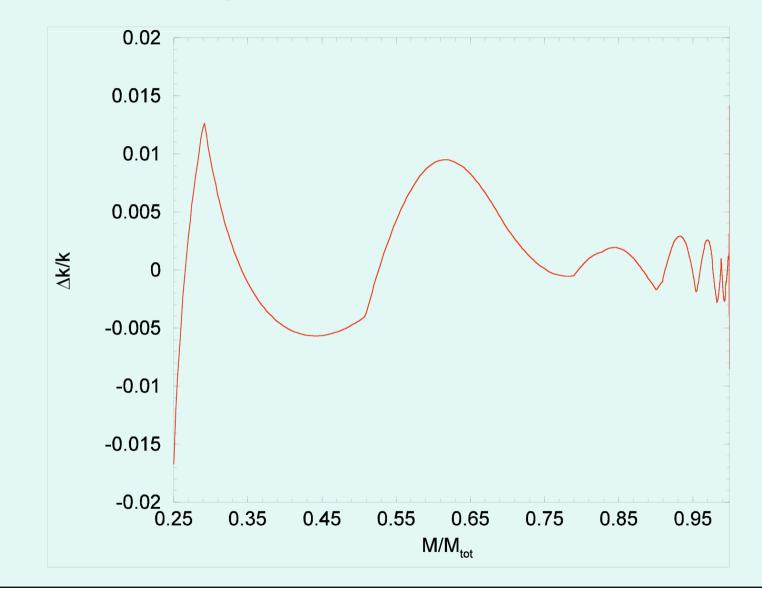




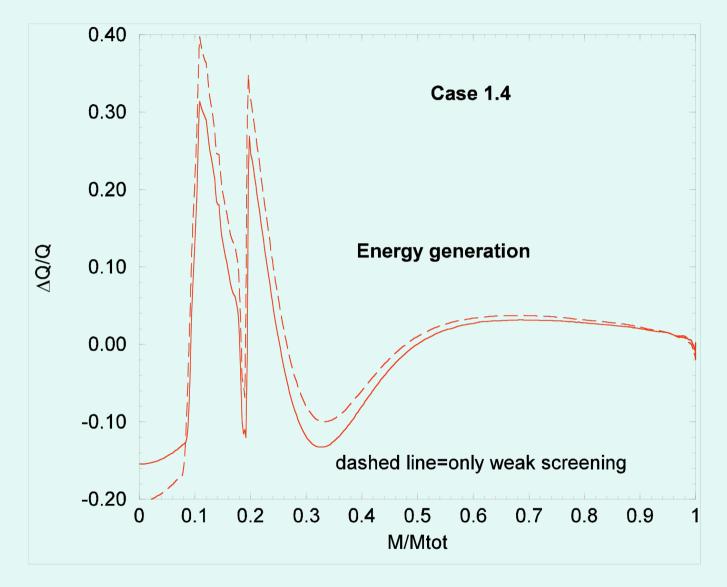
Relative difference for the opacity values



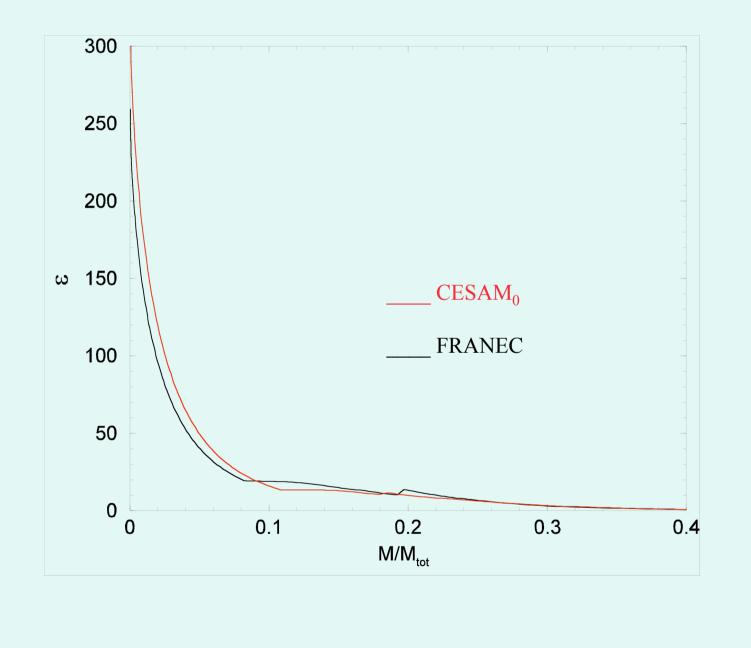
Relative difference between the CESAM_0 opacity and the opacity calculated with our subroutine for T, ρ , X taken from the CESAM_0 model

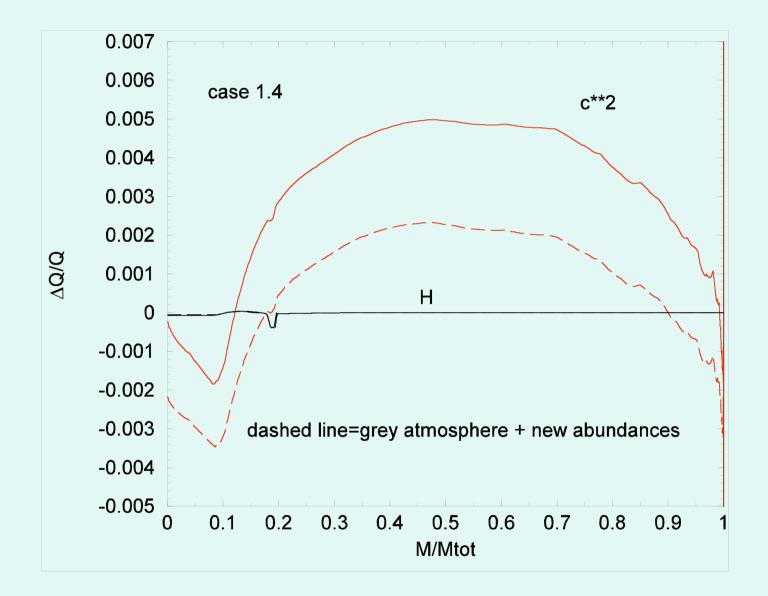


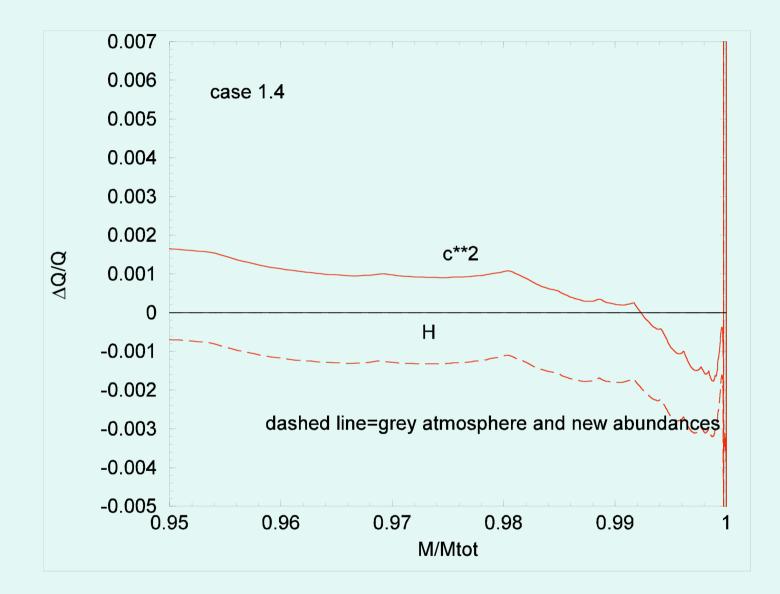
Relative difference for the energy generation

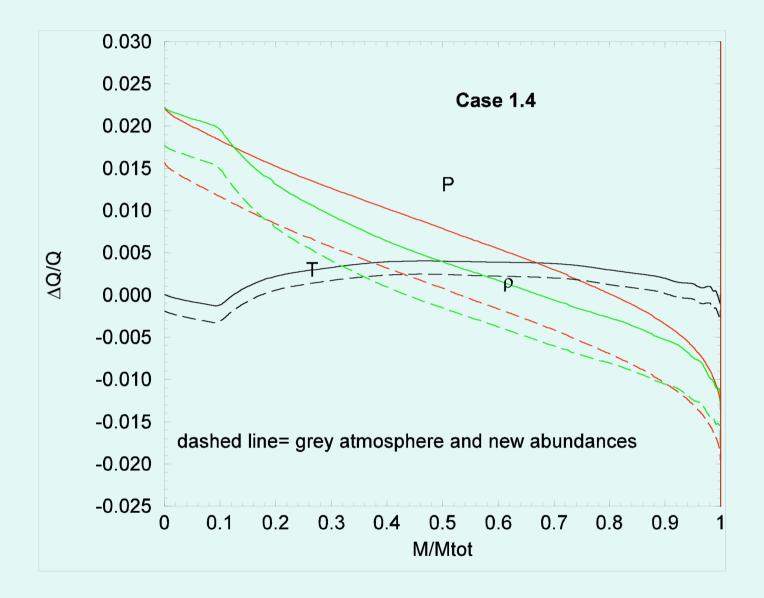


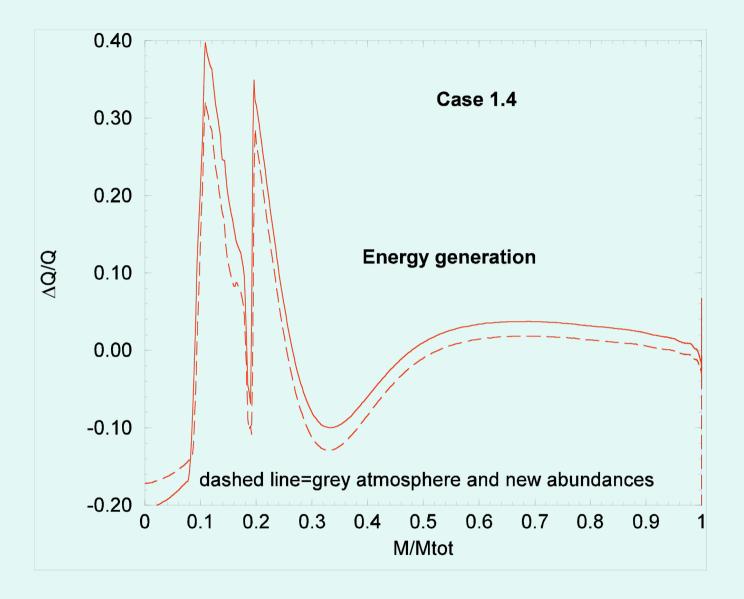
Energy generation profiles











2) Deeper comparison of the case 1.1

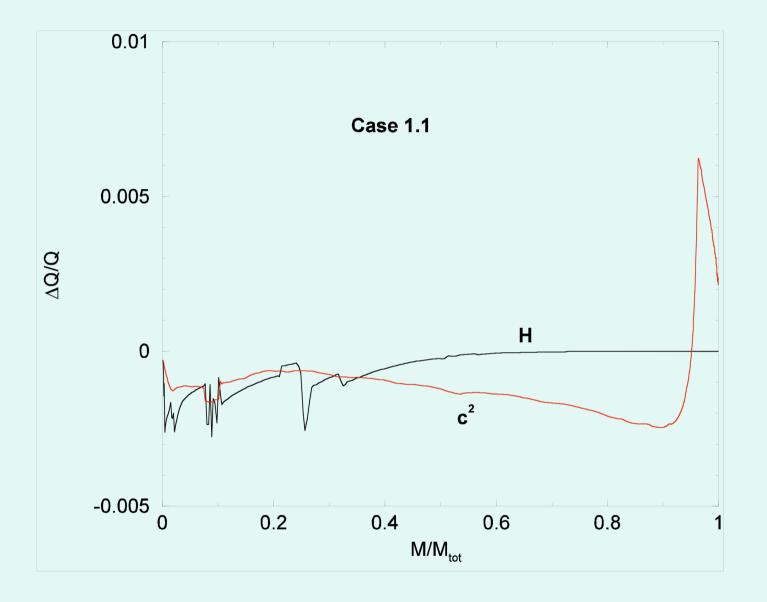
Parameters:

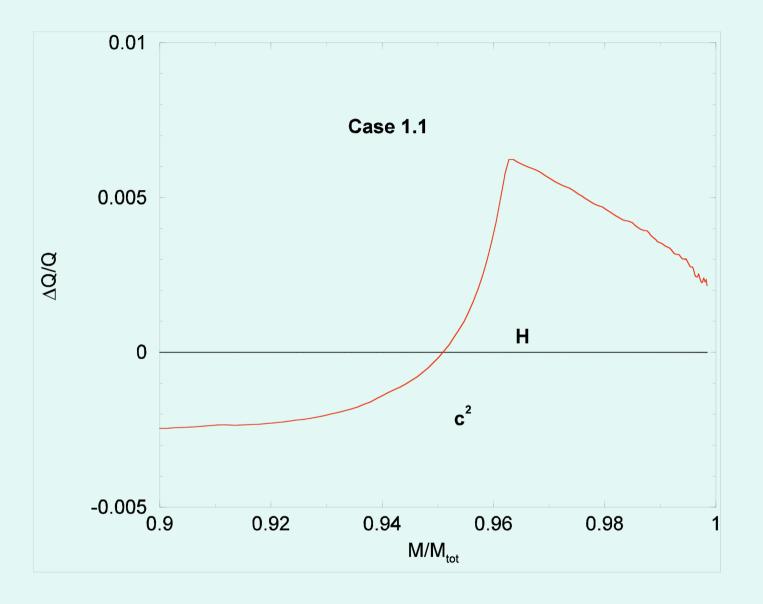
 $M/M_0=0.9$ $X_0=0.70$ $Y_0=0.28$ $Z_0=0.02$ $I_{MLT}/H_p=1.6$

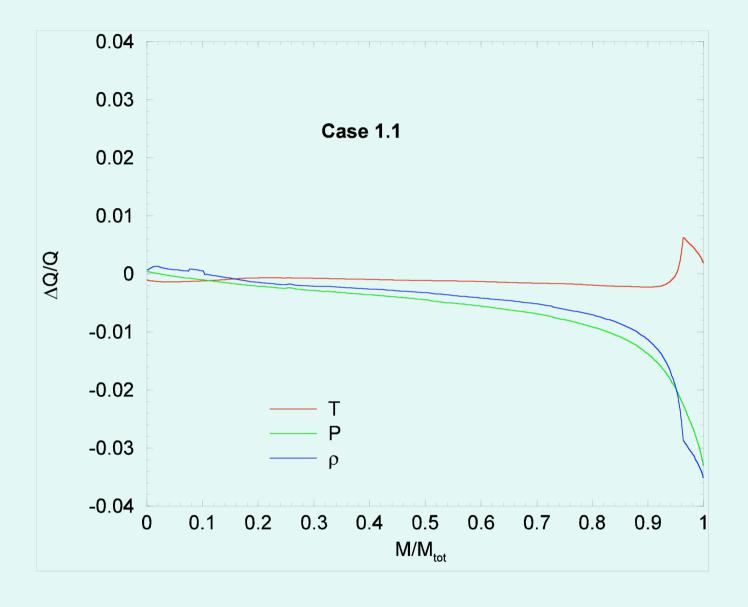
Target model: $X_c=0.35$ (MS)

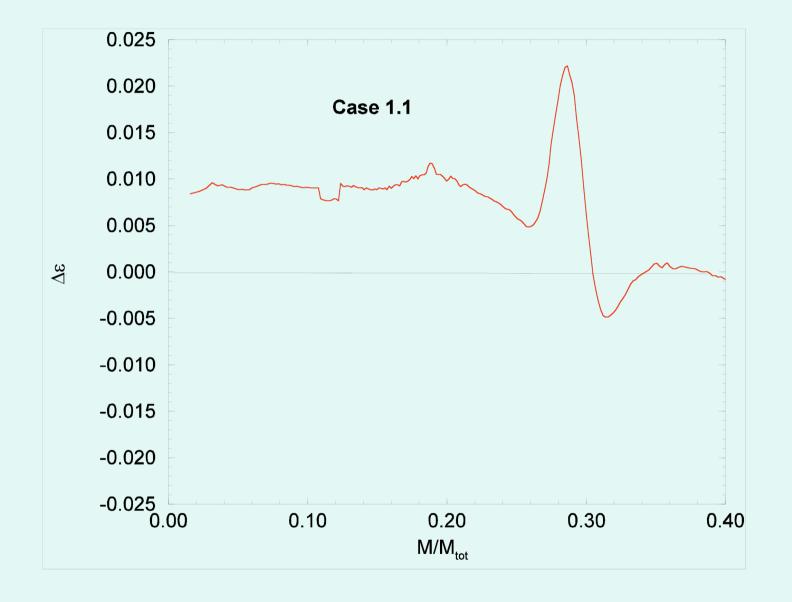
Global parameters case 1.1:

	CESAMo	FRANEC	Δ
•Age [Myr]	6782	6839	0.008
·R/R _o	0.8916	0.8997	0.009
·L/Lo	0.6262	0.6273	0.002
۰T _{eff} [K]	5443	5421	-0.004
•T _c [10 ⁷ K]	1.448	1.446	-0.001
• $ ho_{c}$ [g/cm ³]	150.9	151.0	0.0007
•X _c	0.3501	0.3499	-0.0006
•R _{env} /R	0.6958	0.6972	0.002









Summary:

•For the case 1.4 good agreement for C^2 , X, P, T, ρ but relevant discrepancies for the energy production profile, extension of the convective core, total luminosity -> more tests are needed

•For the 1.1 case good agreement of all the quantities

Future developments:

•All the others task1 models will be recalculated with weak screening, grey atmosphere, new element abundances and no conductive opacity

•The density derivative needed for the pulsation a(i) quantities will be computed analytically and the results will be compared with the ones obtained by other codes